

Small-scale thermal upwellings under the Northern East African Rift from *S* travel-time tomography

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Key Points

- (1) *S* -wave tomography shows multiple upper-mantle upwellings below the Northern East-African Rift as previously proposed using *P* tomography
- (2) Both *S* and *P* anomalies are consistent with a dominantly thermal signature of 100-200 degrees excess temperature
- (3) The complex transition zone structure may well connect to lower-mantle roots revealed in large-scale tomography

Abstract. There is a long-standing debate over how many and what types of plumes underlie the East African rift and whether they do or do not drive its extension and consequent magmatism and seismicity. Here we present a new tomographic study of relative teleseismic S and SKS residuals that expands the resolution from previous regional studies below the northern East African Rift to image structure from the surface to the base of the transition zone. The images reveal two low-velocity clusters, below Afar and west of the Main Ethiopian Rift, that extend throughout the upper mantle and comprise several smaller-scale (about 100-km diameter) low-velocity features. These structures support those of our recent P tomographic study below the region. The relative magnitude of S to P residuals is around 3.5, which is consistent with a predominantly thermal nature of the anomalies. The S - and P -velocity anomalies in the low-velocity clusters can be explained by similar excess temperatures in the range of 100-200°C, consistent with temperatures inferred from other seismic, geochemical and petrological studies. Somewhat stronger V_S anomalies below Afar than west of the MER may include an expression of volatiles and/or melt in this region. These results, together with a comparison with previous larger scale tomographic models, indicate that these structures are likely small-scale upwellings with mild excess temperatures, rising from a regional thermal boundary layer at the base of the upper mantle.

1. Introduction

East African extension and its accompanying volcanic and tectonic activity (Fig. 1a) are commonly associated with one or more mantle plumes rising below the region [e.g. *Burke*, 1996; *Ebinger and Sleep*, 1998]. Indeed, global and continental-scale models have consistently found strong low-velocity regions below the continent, both in the shallow and deep mantle [e.g. *Fishwick*, 2010; *Ritsema et al.*, 1999]. There is however much debate about whether the deeper structure connects with that of the upper mantle.

Some groups have advocated the existence of a single strongly tilted upwelling rooted in the African Large Low Shear-Velocity Province and rising to fill most of the upper mantle from Tanzania to the Red Sea. This is consistent with the pervasive low velocities in the region [*Hansen et al.*, 2012; *Ritsema et al.*, 1999], broad-scale uplift [*Daradich et al.*, 2003], and a consistent pattern of along-rift directions of fast seismic anisotropy [e.g. *Gao et al.*, 2010; *Hammond et al.*, 2014; *Kendall et al.*, 2005; *Montagner et al.*, 2007]. However, others have proposed the existence of at least two separate branches from the lower into the upper mantle, based on more detailed tomographic imaging [*Chang and Van der Lee*, 2011; *Debayle et al.*, 2001; *Koulakov*, 2007; *Montelli et al.*, 2004b], variations in lava chemistry [*Furman et al.*, 2004; *George et al.*, 1998; *Pik et al.*, 2006; *Rogers et al.*, 2000] as well as modelling of the geoid and the evolution of magmatism [*Lin et al.*, 2005]. A recent P tomographic study we performed below the northern East African Rift [*Civiero et al.*, 2015] combined regional seismic datasets across the region to obtain resolution to the base of the transition zone and revealed even smaller structures (scales ~ 100 km) extending across the depth of the upper mantle. We proposed these to be small-scale upwellings rising from a zone of (ponded) hot material at the base of the upper mantle. Such small-scale structures are consistent with those imaged by previous tomographic studies with resolution down to about 400 km depth [*Bastow et al.*, 2005; *Bastow et al.*, 2008; *Hammond et al.*, 2013], and also with proposed

interpretations of strongly variable isotopic characteristics of lava along the rift [Furman *et al.*, 2006; Meshesha and Shinjo, 2008].

Here we present a complementary tomographic inversion of relative S and SKS travel times to (a) assess robustness of the structure revealed by our P tomography [Civiero *et al.*, 2015], which we refer to as NEAR-P15, and (b) more importantly, to further constrain the thermal and/or chemical nature of the structures.

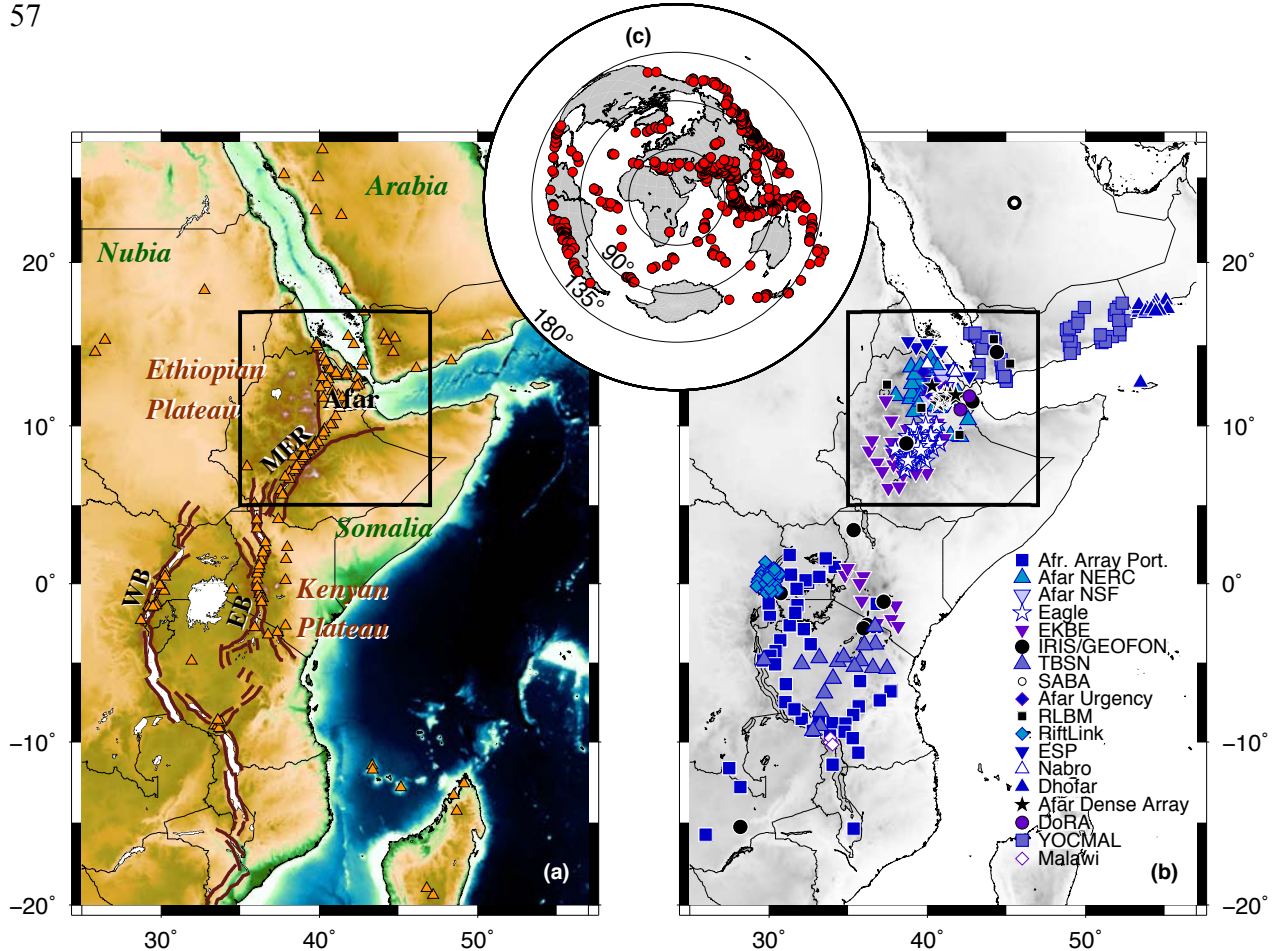


Figure 1. (a) Tectonic map of the East African Rift, showing the presence of two topographic domes, the main faults of the rift system (in brown), with the Main Ethiopian Rift (MER) cutting through the Ethiopian Plateau and a Western (WB) and Eastern Rift Branch (EB) around the Kenyan Plateau. Holocene volcanoes (in orange) largely concentrate along the rift zone (Smithsonian Volcanism Program 2013). The black rectangle delineates the area of interpretation. (b) Distribution of all stations used in this study, symbol-coded according to their network. Network and station information can be found in supporting information [Table S1](#). (c) Distribution of earthquake sources used for the S - SKS tomography. Source-receiver distances shown by the circles on the inset are from the centre of the black rectangle in (a) and (b).

2. Method

2.1. Data

This study uses broadband recordings of teleseismic S - and SKS -wave phases from 379 stations across eastern Africa, from Malawi to Eritrea, and the Arabian Peninsula (i.e., all those shown in [Fig. 1b](#)). The seismic stations belong to 16

multinational projects and overlap spatially and/or temporally in our region of interest (see supplementary material). There has been some confusion in the literature about the resolving capability of the different generations of tomographic models in the north East-Africa rift (e.g., Reed et al., 2016). Early studies used stations in and around the Main Ethiopian Rift, resolving P - and S -wave velocities in the top 300-400 km (Bastow et al., 2005, Benoit et al., 2006, Bastow et al., 2008). The data set was extended into Afar and Eritrea by Hammond et al., 2013 who focussed their interpretation on the top 400 km in relation to rifting in Afar. However, due to the increased aperture of their seismic array resolution, particularly the S -wave model, due to the inclusion of significant SKS data, was good to depths of ~ 600 km. As a result, the Hammond et al., 2013 model was used to compare against images of transition-zone discontinuity structure in Thompson et al., 2015. The most recent models of Civiero et al., 2015 and this study have extended the data significantly further, including data from Saudi Arabia to Malawi. As shown in Civiero et al., 2015 and Fig. S8 this substantially increases crossing rays in and around the transition zone and thus allows us to confidently interpret seismic velocity models to the base of the transition zone.

We pick arrival times from 590 earthquakes of magnitude $m_b \geq 5.5$, ranging in epicentral distance from 30° to 130° (30° - 90° for S waves, 90° - 130° for SKS waves). The azimuthal and distance coverage provided by the selected earthquakes is displayed in Fig. 1c. After applying a 0.04-0.15 Hz band-pass Butterworth filter, S picking was performed on the transverse component, to minimise the effect of P , P - S and S - P conversions, while SKS picking was done on the radial component, as it is dominated by SV energy. Final picks and relative arrival times were determined using the multichannel cross-correlation method (MCCC) of *VanDecar and Crosson* [1990], in a 12 s window around the initial pick. We deleted all picks with cross-correlation coefficients < 0.80 from our analysis. In total, we retain 16569 travel-time picks, divided into 8730 and 7839 S - and SKS -wave phases, respectively. The mean RMS uncertainty in the delay times, based on the number of picks rather than on cross-correlation pairs for each event [Tilmann et al., 2001], is 0.38s with a standard deviation of 31 s (S -waves mean RMS uncertainty = 0.34s with standard deviation = 0.28s and SKS -waves mean RMS uncertainty = 0.43s with standard deviation = 0.33s). The range of S - and SKS - travel-time residuals is about ± 10 seconds. To reduce the effect of a few higher outliers, residuals larger than 3.0 standard deviations are downweighted iteratively in the inversion [Huber, 1981]; 827 rays end up downweighted, i.e. about 5 % of the data.

2.2. Tomographic inversion

The data are inverted using the teleseismic travel-time inversion method of *VanDecar et al.* [1995], with the same scheme as in our previous P -wave tomographic analysis [Civiero et al., 2015], which has been used and described in numerous previous studies [e.g. Hammond et al., 2013; Ritsema et al., 1998; Schimmel et al., 2003; Wolfe et al., 1997; Wolfe et al., 2009]. We use ray theory and perform a linear inversion; both assumptions should not have a substantial effect on the shape of the anomalies but may somewhat affect amplitudes [e.g. Bastow et al., 2005; Montelli et al., 2004a; Peter et al., 2009]. There is insufficient information on crustal structure below the whole network to correct the data for station statics, so we solve for station (and event) corrections as part of the inversion. We use exactly the same parameterisation as in NEAR-P15, with an inner grid below the stations of $0.5^\circ/0.4^\circ$

node spacing in latitude/longitude and 50 km in depth, extending down to 1500 km, and an outer grid with 1° node spacing and 100 km depth spacing extending to 28°N-25.40°S in latitude, 25°E-57.20°E in longitude and 2000 km in depth, to avoid mapping outside structure into our region of interest.

We regularise the model by suppressing spatial and curvature gradients (smoothing and flattening), as well as, in part of the cases, by damping towards a 3D reference structure. We first investigate the trade-off between the RMS residual reduction and RMS model roughness for different combinations of smoothing and flattening to choose a preferred model that fits the data well, i.e. within a residual reduction that can be justified by the estimated noise level. We obtain the same preferred regularisation parameters for our final model (which we will call NEAR-S16), as for NEAR-P15, although they were determined independently. Preferred flattening and smoothing factors are 4800 and 153600, respectively (Fig. 2).

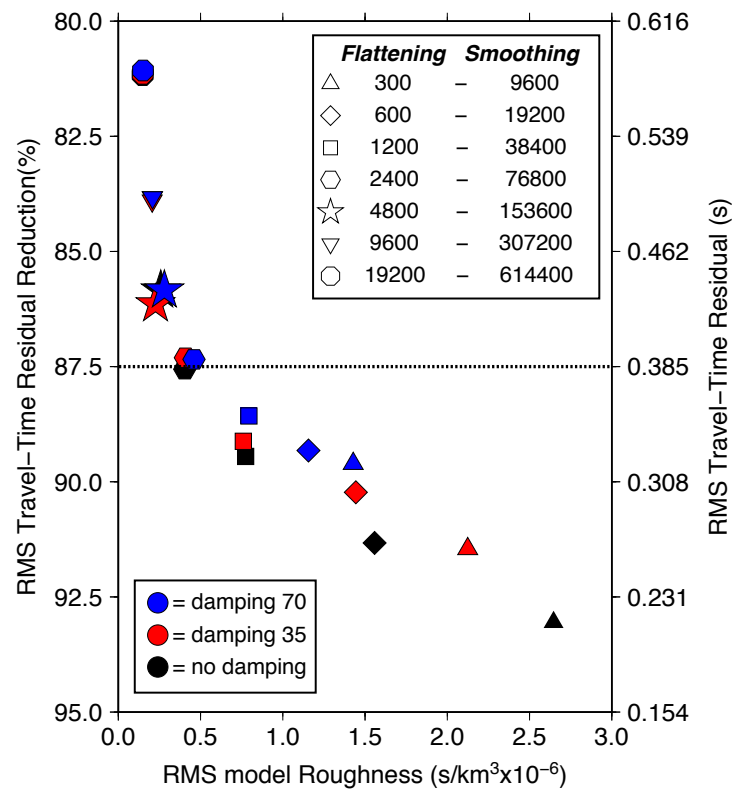


Figure 2. Trade-off between *S/SKS*-velocity model roughness and data fit for inversions with different degrees of flattening and smoothing (symbols) and three different degrees of damping toward the surface-wave constrained initial structure (colours). The dashed line represents our estimate of travel-time residual error (see text for details). The regularization parameters for our preferred model are 4800 for flattening and 153600 for smoothing and a damping of 35 (red star).

In most previous travel-time inversions for the region, a 1-D starting model was used (i.e., zero anomaly, as the relative travel-time method has no sensitivity to 1-D structure) and no explicit damping to it was performed [Bastow *et al.*, 2005; Bastow *et al.*, 2008; Hammond *et al.*, 2013]. However, teleseismic travel-time tomography provides relatively poor resolution of the crustal and uppermost mantle structures above ~200 km depth because of the paucity of crossing ray paths in that depth range. Regional surface-wave data can resolve this shallow part of the upper mantle better.

Therefore, as for our NEAR-P15 model, we damp the inversion using the regional surface-wave model of *Fishwick* [2010] [F2010] as our starting model above 350 km depth (shown in supplementary Fig. S1). The surface wave model is estimated to reach a vertical resolution of ~50 km and lateral resolution between 150 and 500 km. In the area of interest, given the large number of stations and path coverage, we are using one of the best-resolved parts of the model. Where required by the travel-time data, our inversion will add smaller-scale structure to this long-wavelength starting model. Below 350 km, we damp to a 1D (zero-anomaly) reference. The benefits and effects of such damping are discussed in detail in *Civiero et al.* [2015].

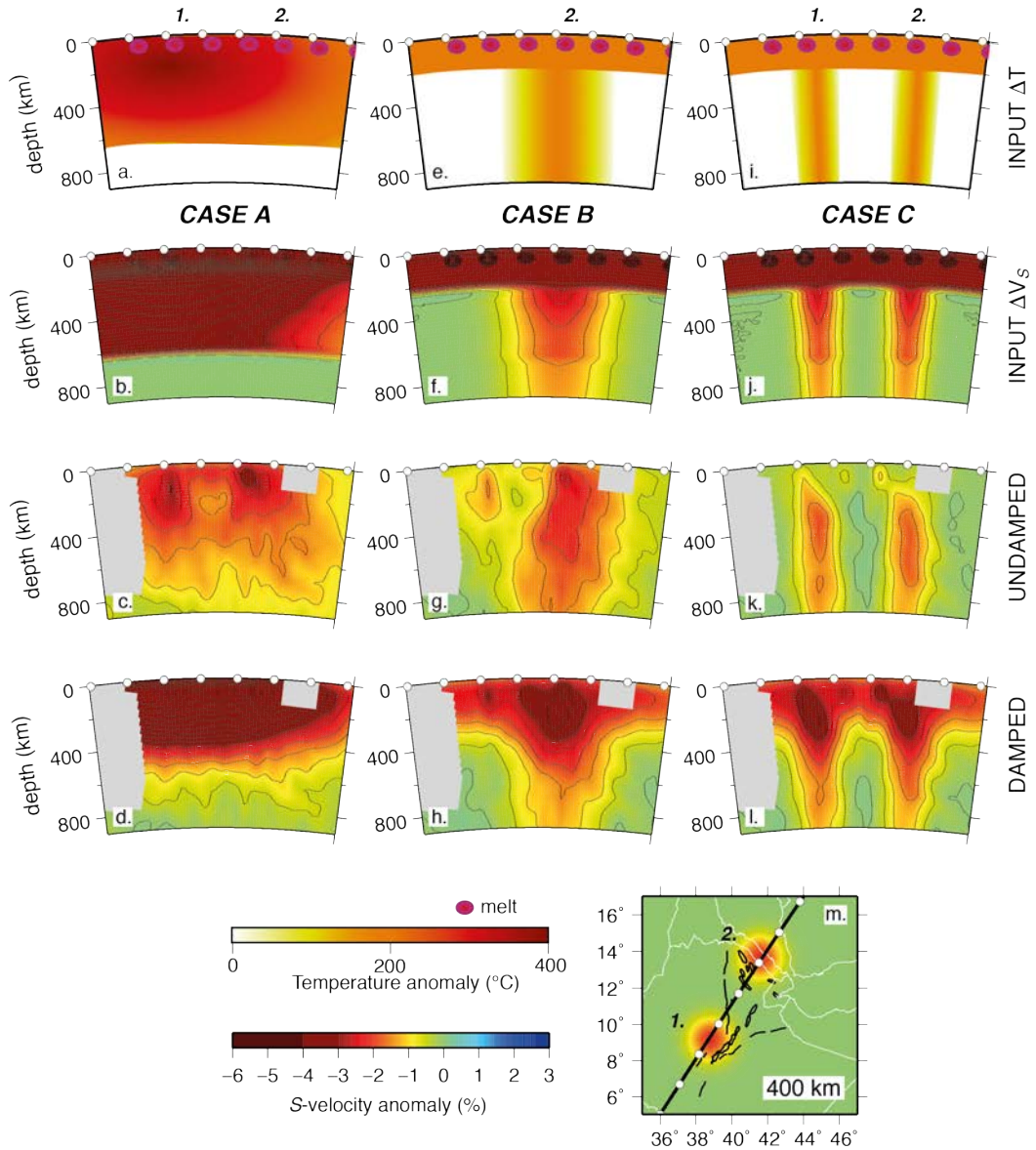
We test a range of damping factors from 0 (no damping) to a fairly strong factor of 70 (where the shallow part of the model is largely the same as F2010). As was found for the *P*-wave models, the RMS residual reduction does not change significantly with changes in damping (for our preferred flattening and smoothing factors of 4800 and 153600, the residual reduction is 85.8, 86.1, 85.8%, for a damping of 0, 35 and 70, respectively). However, the heavily damped case (damping = 70) has stronger shallow anomalies and weaker (but still required) deeper structure while the undamped case (damping = 0), which is fully determined by the travel times, shows the same spatial distribution of anomalies, but has weaker shallower and stronger deeper anomalies (Suppl. Fig. S3, S4).

3. Tomographic model

3.1. Resolution

We performed a range of resolution tests to assess the quality of our retrieved tomographic models. Standard checkerboard tests are included in the supplementary material (Fig. S9 and S10), as well as several tests for how much smearing occurs from structures above and below the transition zone into the transition zone (Fig. S11). In addition, we run a set of tests for the three alternative plume hypotheses (Fig. 3): (A) a large-scale plume head filling all of the upper mantle, (B) a single plume (with assigned width of 380 km, defined as the distance where the anomaly amplitude is 20% of that in the centre of the synthetic plume structure), (C) two narrow vertical plumes (with widths of 150 km). For case B and C a hot layer is added at the top of the input structures to mimic plume material spread in the rift zone below the plate. As a starting model for the damped plume resolution tests, we used the corresponding synthetic input models, filtered to mimic the effect of the surface wave resolution. The filter applies a variable vertical averaging (lower vertical resolution as absolute depth increases) and a variable Gaussian spatial filter (lower spatial resolution as depth increases).

The checkerboard tests illustrate that we achieve quite good resolution for anomalies of 125 km and 250 km diameter through most of the transition zone (Suppl. Fig. S9, S10). At shallow depths (around 100 km), the resolved area is confined to directly below the stations, and from about 300 km depth covers much of the region of interest shown in Fig. 1. At the edges of the model, the checkers become more smeared with depth. For 125-km sized features, resolution becomes poor below about 600 km depth for most of the area (Fig. S10). Larger features, with diameters of at least 250 km, are resolved over a wider region and through the transition zone, down to 700-800 km depth (Suppl. Fig. S9).



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207 **Figure 3.** Three synthetic tests to examine the resolving power of our tomographic inversion
 208 for three previously proposed mantle upwelling structures: Case A - superplume (a-d), Case
 209 B - single plume (e-h) and Case C - multiple plumes (i-l) scenarios. Vertical cross sections
 210 through the input and output models are shown in panels a-l. The location of the cross section
 211 is shown on a depth slice through the two-plume velocity input in panel m. Input models are
 212 defined in terms of temperature anomalies (a,e,i) and then multiplied by the isomorphic
 213 derivative from Suppl. Fig. S12 to obtain V_s anomalies (panel b,f,h). In all cases, a set of 4%
 214 V_s checker anomalies below the rift is added to the plume structures above 100 km depth to
 215 mimic lateral variability due to lithospheric topography and melt concentration. (a-d) Case
 216 A: superplume represented by a flat ellipsoid with a maximum excess T of 400°C at 120 km
 217 depth, and a Gaussian fall off with depth leading to 0°C anomalies at 660 km; (e-h) Case B:
 218 single larger upwelling modelled as a vertical cylinder positioned beneath Afar with
 219 maximum 200°C excess temperature that is constant with depth and Gaussian variation
 220 laterally over a 380 km width (defined as the distance to 20% of the maximum amplitude),
 221 plus a layer of 200°C excess temperature above 200 km depth; (i-l) Case C: Two smaller
 222 upwellings represented by vertical cylinders, positioned beneath Afar and MER, again
 223 assuming a maximum 200°C T anomaly that is constant with depth, a Gaussian variation
 224 laterally over a 150 km width and a 200°C hot layer above 200 km depth. (c,g,k) Recovered

models using the same parameters as in the data inversion, but with no damping. *(d,h,l)* recovered models using the same parameters as in the data inversion, including damping (factor = 35). The starting model to which the inversion is damped is a version of the 3-D synthetic model down to a depth of 350 km smoothed to mimic spatial resolution of F2010. Regions with less than 5 rays per node are shaded grey. The spacing between the contours is 0.25%. White points indicate the distance every 2 degrees. Profiles through the centre of the plumes showing the input and retrieved anomalies are shown in Suppl. Fig. S15. These tests illustrate that without damping most shallow structure is not recovered. Damped inversions yield a better representation of the vertical continuity of the plume structures.

The plume tests in Fig. 3 illustrate the importance of using a-priori constraints for the shallow structure in the starting model and damping towards these structures. Without such constraints, the expected laterally widespread low velocity structures above 200 km depth can not be retrieved. While undamped inversions emphasise lateral variations in velocity structures, which in the case of models B and C means the plume tails, in the damped inversions the geometry of the imaged structures is much closer to that of the input structures. Note that due to the decrease in thermal sensitivity of V_s with depth (from about 3%/100K at 100 km depth to <1%/100K at 600 km depth, Suppl. Fig. S12) and the expectation that temperatures of adiabatic upwellings only change by a few 10s of degrees over the depth of the upper mantle, we would expect plume anomalies to decrease with depth. Thus tomographic images of upper mantle plumes are not expected to have simple cylindrical morphologies.

3.2. *S-velocity structure*

Our preferred *S-SKS* model (with intermediate damping) is shown in Fig. 4. Undamped and strongly damped models are included as Suppl. Fig. S3 and S4. The colour scale used is the same in all panels to illustrate how anomaly amplitudes vary with depth. The scale best illustrates the transition zone anomalies and saturates at depths from 300 km upwards for which the structures have been analysed in detail in previous studies [Bastow *et al.*, 2005; Bastow *et al.*, 2008; Hammond *et al.*, 2013]. The damped models in particular emphasise that the strongest low velocities are shallow, where they are widespread. Deep structures persist at all degrees of damping and are hence required by the data. Also, the synthetic test for a large anomaly confined to the top 400 km shows limited smearing to deeper depths (Suppl. Fig. S11), illustrating that the transition zone anomalies are not a result of smearing.

Station terms are mostly negative below the oceanic Red Sea and generally positive beneath the African continent. Some of the negative station corrections under Arabia and the Afar Depression may in effect be a compensation for lateral smearing of the low velocity from northeastern Africa in the surface-wave model (compare with the undamped case in Suppl. Fig. S3 where these stations have positive corrections). Overall, the station correction trend reflects crustal structure [e.g. Hammond *et al.*, 2011].

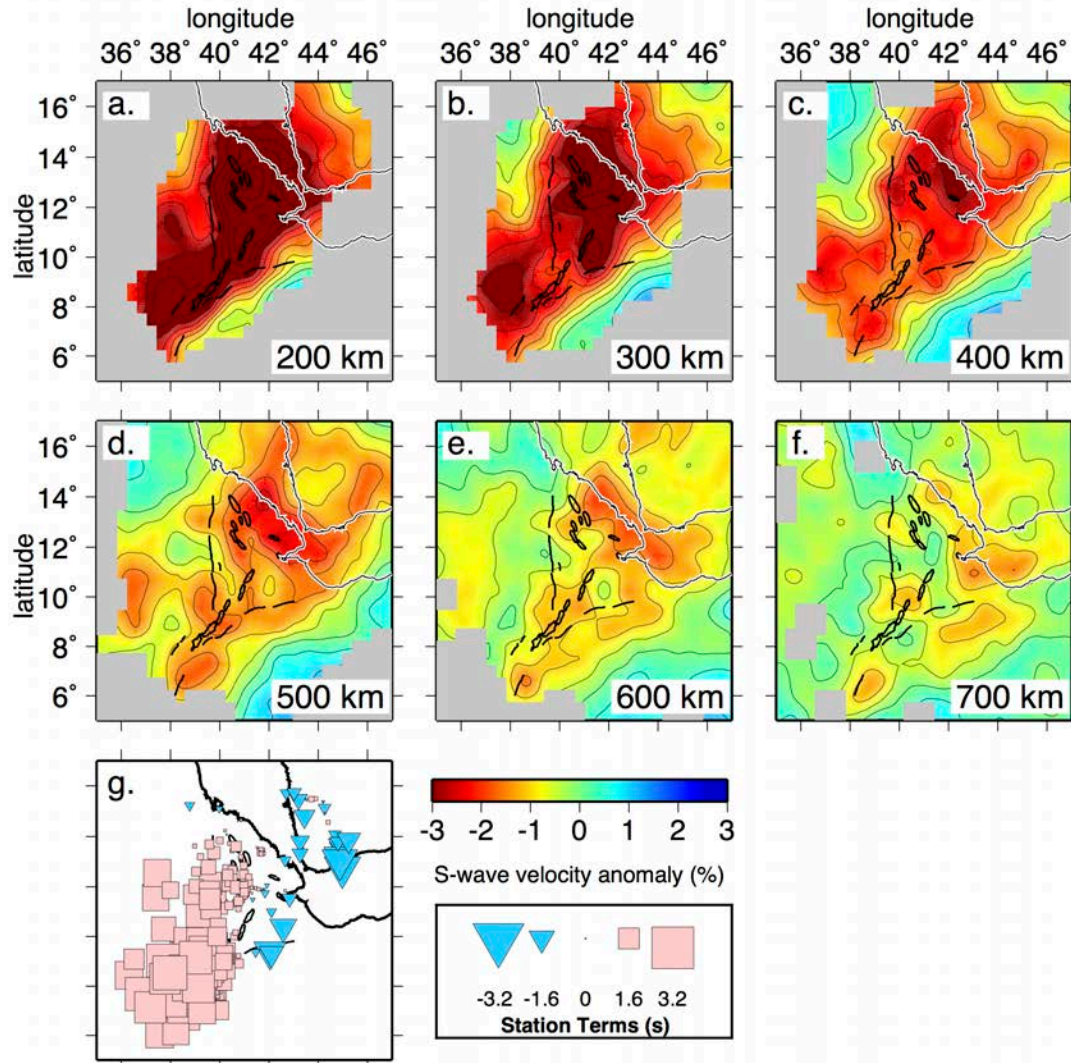


Figure 4. Depth slices through our *S-SKS* preferred *intermediate-damped* tomographic model (flattening=4800, smoothing=153600, damping=35), at depths between 200 and 700 km. Regions with less than 5 rays per node are shaded gray (a version where regions with less than 10 rays per node are shaded can be found in Fig. S2). The spacing between the contours is 0.5%. Black lines delineate the major border faults and magmatic zones bounding the Afar Depression and black over white lines show coastlines. Triangular and square symbols in panel g represent the sign and magnitude of the station static terms. Compare with Suppl. Fig. S5 to see the first-order similarities in structure between the independently inverted *S/SKS*- and *P*- velocity models. Compared to the undamped model (Suppl. Fig. S3), the shallow mantle anomalies are enhanced and more extensive. Otherwise, similar features, including the low velocities below Afar and west of the MER, appear.

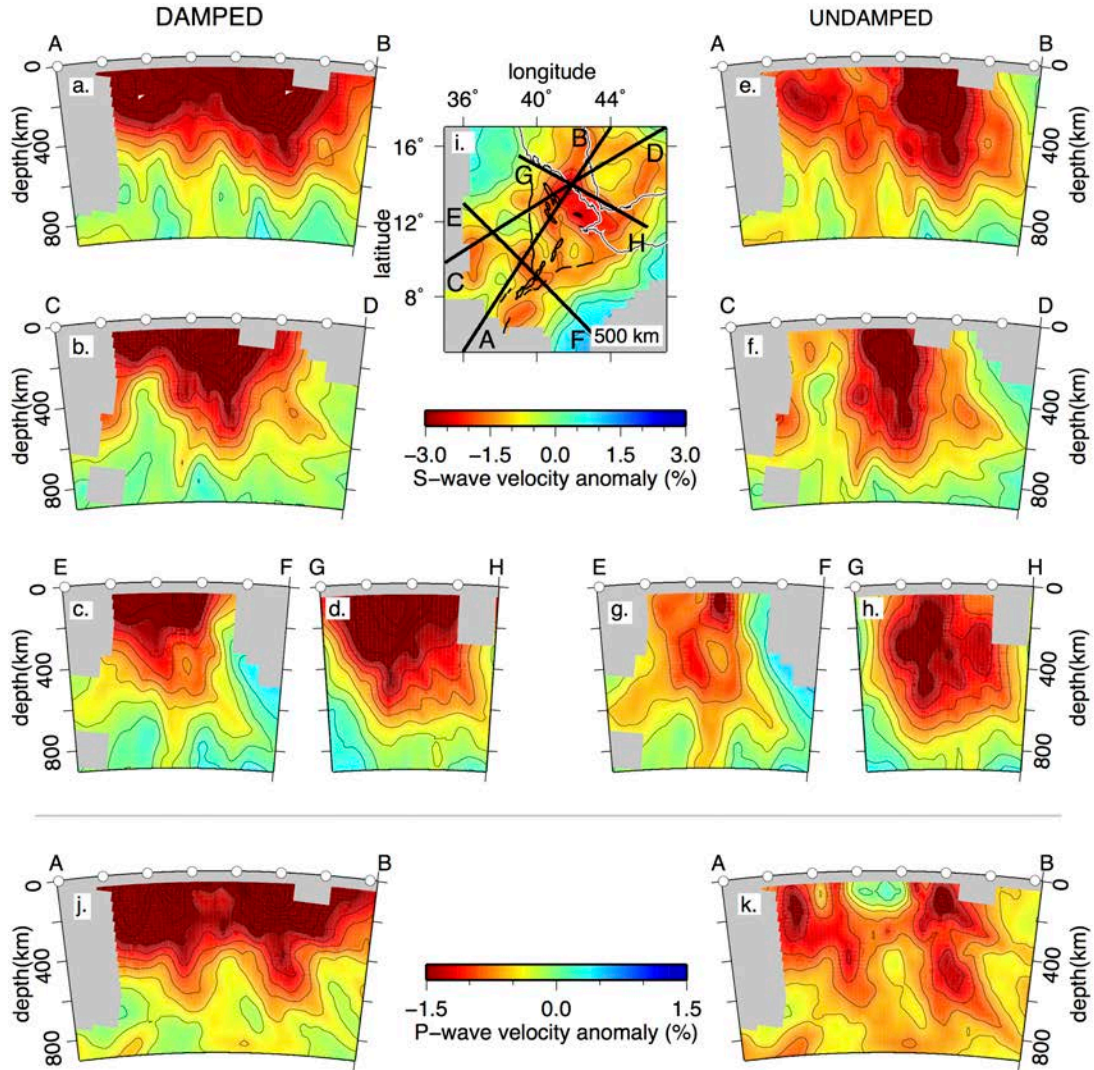


Figure 5. Vertical cross sections through the S-SKS model (a-i) and comparison with a section through NEAR-P15 (j,k). Regions with less than 5 rays per node are shaded gray. The spacing between the contours is 0.5% for the S model, 0.25% for the P model. White circles along the top of the vertical profiles mark the distance every 2° . (a–d) damped (damping=35) and (e–f) undamped (damping=0) S-SKS models (both have flattening=4800, smoothing=153600). The location of the cross sections (black lines) is shown in the 500 km depth slice through the damped model (i). Cross section A-B, C-D, E-F and G-H have the same orientations as the profiles through the P-velocity models in (j,k) and Suppl. Fig. S6. The undamped models (e–h) emphasize the stronger amplitude of the low-velocity cluster below Afar compared to that beneath MER; the damped models (a–d) show that the low-velocity anomalies extend from the surface to at least the bottom of the transition zone. The features in S model are similar to those seen in the P model (j,k).

The mantle structure shows the same first-order features as NEAR-P15 [Civiero *et al.*, 2015]. The shallow upper mantle, down to ~ 100 km, reveals a broad low-velocity layer with the strongest anomalies following the rift morphology ($-4.5\% < \delta V_S < -3.0\%$), similar to other previous tomographic studies for this region [e.g. Bastow *et al.*, 2008; Fishwick, 2010; Hammond *et al.*, 2013].

At larger depths, the low-velocity structure splits into several separate anomalies, that persist down to at least ~ 650 km depth. In NEAR-P15, these deeper anomalies clearly form two clusters, each about ~ 400 km in diameter, beneath Afar and west of

the MER (Fig. S5), correlating with active volcanism in the Afar region and volcanism along and off-axis of the MER. In the *S-SKS* model, the same two clusters are also clear from about 200 to 400 km depth. In the transition zone, the Afar cluster is more strongly defined than the one below the MER, which breaks up into several smaller-scale features. Note that, in NEAR-P15, the clusters also contained additional small-scale structure in the transition zone (Fig. S5, S6).

The amplitude of the low-velocity features in the intermediately damped case decreases with depth from $\sim -4\%$ in the uppermost mantle to $\sim -0.5\%$ below the transition zone. The deepest resolution is too limited to determine how the transition zone structure relates with that below 600-700 km depth.

Complexity in the structure may be to some extent influenced by anisotropy. However, we have quite good azimuthal coverage (as shown in supplementary Fig. S8), where the core of the region is traversed by rays spanning an azimuthal range of 270-360 degrees, which should aid in averaging out azimuthal anisotropic effects. And although we combine SH and SV data, we get images that are quite consistent with those of the *P* data, so if radial anisotropy exists it does not appear to map significantly into the *S* structures we resolve. Nonetheless, this would be an aspect that warrants future work.

Vertical cross sections through the models (Fig. 5) further illustrate the vertical continuity of the two clusters from the shallow mantle through the transition zone, again consistent with the *P*-wave structure. However, different from the *P*-velocity structure (Suppl. Fig. S5, S6), the anomalies below Afar are stronger in amplitude and more continuous than the anomalies west of the MER (see profiles E-F and G-H in Fig. 5).

4. Physical interpretation:

4.1. Ratio of *P* and *S* velocity anomalies

Several studies have used the ratio of relative changes in shear- and compressional-wave velocities, defined as $R_{S,P} (= d\ln V_S / d\ln V_P = [dV_S/V_S] / [dV_P/V_P])$, as diagnostic of the cause of the seismic anomalies [e.g. Karato and Karki, 2001; Masters *et al.*, 2000; Robertson and Woodhouse, 1996; Simmons *et al.*, 2009]. The expected $R_{S,P}$ for purely thermal variations spans a range from 1.3 (for low values of temperature, with a weak anelasticity effect) to 2.2 (at high temperatures, where the anelasticity effect is strong) [Cammarano *et al.*, 2003; Goes *et al.*, 2000]. Where water is present, it may enhance the anelastic sensitivity to temperature, by lowering the melting temperature [Karato and Jung, 1998] and/or form hydrated minerals that usually have lower velocity than average mantle minerals [e.g. Abers and Hacker, 2016; Angel *et al.*, 2001]. The anelastic effects of fluids can increase $R_{S,P}$ to up to about 2.3; hydrous minerals, if present in large enough quantities possibly somewhat more [e.g. Goes *et al.*, 2000; Hacker and Abers, 2012].

For most plausible large-scale compositions (i.e., more or less melt-depleted forms of peridotite, or alternative chondritic compositions), the effect of composition on V_P and V_S above the transition zone is small compared to that of temperature [Afonso *et al.*, 2010; Cammarano *et al.*, 2003; Goes *et al.*, 2000; Worthington *et al.*, 2013]. In the transition zone, the sensitivity of velocity to changes in composition is complicated by the different depths, Clapeyron slopes and widths of the phase transitions expected for different bulk compositions [Cammarano *et al.*, 2003; Xu *et al.*

al., 2008]. Overall, for most upper-mantle compositions and depths, the sensitivity of $R_{S,P}$ to temperature and composition is likely not different enough to distinguish the two factors.

The largest effects on $R_{S,P}$ are expected from partial melt [e.g. *Faul et al.*, 2004; *Hammond and Kendall*, 2016; *Hammond and Humphreys*, 2000; *Schmeling*, 1985] with values that can range from around 1.6 for ellipsoidal melt inclusions, to 2.2 for films in a geometry taken from natural samples [*Hammond and Humphreys*, 2000], to up to 4.0 for aligned film and layer geometries [*Takei*, 2002], which would also lead to strong anisotropy.

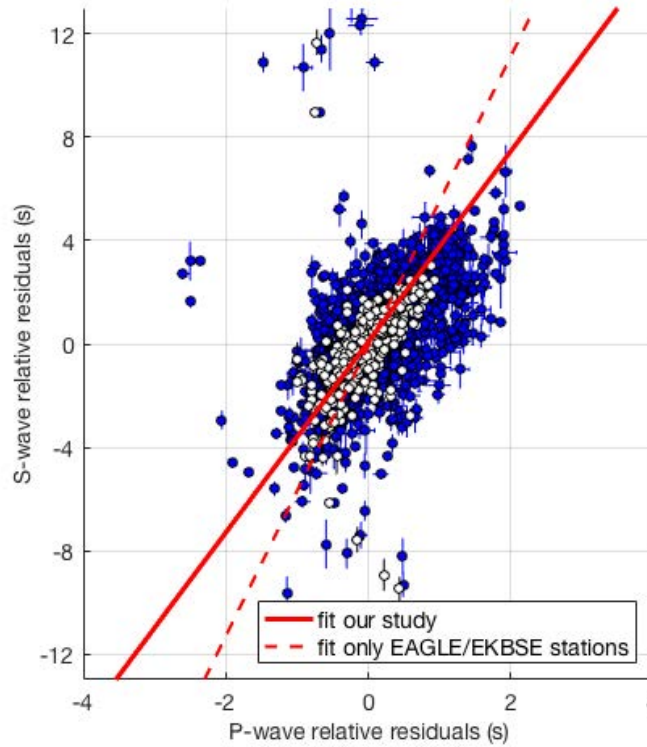


Figure 6. *P*-wave (from *Civiero et al.* 2015) versus *S*-wave (this study) relative arrival-time residuals for all common earthquakes. The solid red line is a least-squares fit including picking errors for all data (blue and white circles) and has a gradient, $\delta_{S,P}$, of 3.69 ± 0.01 for all data corresponding to a ratio $R_{S,P} = \ln V_S / \ln V_P$ of ~ 2.1 , consistent with a dominantly thermal origin of the anomalies. The dashed red line is a least-squares fit including only data close to the MER (white circles) and has a gradient of 5.60 ± 0.03 ($R_{S,P} = \sim 3.2$), previously interpreted as implying the presence of melt in a region around the MER [*Bastow et al.* 2005].

Several studies use the ratio of *P*- and *S*-wave relative arrival-time residuals $\delta_{S,P}$, i.e., using the data rather than the tomographically imaged anomalies [e.g. *Bastow et al.*, 2005; *Gao et al.*, 2004; *Rocha et al.*, 2011]. $\delta_{S,P}$ should be proportional to the ratio of absolute dV_S and dV_P along the chosen station-event pairs. Thus, the slope $\delta_{S,P}$ equals $(V_P/V_S) R_{S,P}$ or about $\sqrt{3}$ times $R_{S,P}$. *Bastow et al.* [2005] perform a $\delta_{S,P}$ analysis below the Main Ethiopian Rift and find a gradient between ~ 5 and ~ 10 , which they attribute to significant fractions of shallow melt ponding beneath the rift, consistent with other geochemical and geophysical results [e.g. *Bastow et al.*, 2005; *Ebinger and Casey*, 2001; *Gao et al.*, 1997; *Hammond et al.*, 2014; *Kendall et al.*, 2005; *Rooney et al.*, 2007]. Other studies have used $\delta_{S,P}$ to argue for chemical

heterogeneity. For example, *Gao et al.* [2004] conclude a $\delta_{S,P}$ of ~ 2.9 for fast anomalies below the Colorado Plateau is on the high side to invoke temperature perturbations. Similarly, *Rocha et al.* [2011] find $\delta_{S,P} > 2.9$ below the São Francisco Craton and interpret it as due to compositional effects. However, from our preceding discussion of the expected range of $R_{S,P}$ values including the effect of anelasticity, $\delta_{S,P}$ up to 3.6-3.8 correspond to $R_{S,P}$ which fall within the thermal range.

To get further insight into the possible physical causes of the seismic anomalies we imaged below the Northern East African Rift, we perform a comparison of the P - and S -SKS-wave relative arrival- time residuals for common events (Fig. 6). Prior to the analysis, we remove all residuals downweighted by the Huber iterations. A straight line fit through our measurements (least-squares, accounting for the estimated picking errors) yields a $\delta_{S,P}$ of 3.69 ± 0.01 which corresponds to an $R_{S,P}$ of ~ 2.1 . This value is towards the higher end of estimates of purely thermal $R_{S,P}$ between ~ 1.3 to 2.2 [*Cammarano et al.*, 2003; *Goes et al.*, 2000], and is as expected, given the likely elevated mantle temperatures (and thus expected higher anelasticity effect) below the region indicated by previous work [*Armitage et al.*, 2015; *Ferguson et al.*, 2013; *Rooney et al.*, 2012].

This result seems contradictory to the previous $\delta_{S,P}$, between ~ 5 to ~ 10 , for stations within the MER [*Bastow et al.*, 2005]. However, when we calculate the $\delta_{S,P}$ only for the subset of stations centred around MER (i.e., more or less corresponding to the area covered by the EAGLE experiment that *Bastow et al.* [2005] analysed), we find a higher slope of 5.60 ± 0.03 corresponding to an $R_{S,P} \sim 3.2$. This indicates that what *Bastow et al.* [2005] interpreted as the signature of melt, is likely local and shallow. Our new travel-time data set covers a significantly longer time span and a larger region of the rift, and thus the effects of local shallow melt signatures are mixed in with deeper mantle thermal signatures.

We also computed $R_{S,P}$ values after the tomographic inversion. These ratios are strongly scattered, ranging from about 0 to 4 (Fig. S13), as they are affected by any differences in spatial resolution of the two models. For this reason, care must be taken interpreting the $R_{S,P}$ distribution. Broadly, the $R_{S,P}$ distribution is peaked in the thermal range (median 1.7 for damping=35). The $R_{S,P}$ in the damped models is more strongly peaked in the thermal range, as it conditioned by our thermally-scaled starting model for P . This does however confirm that a thermal interpretation is compatible with the data. Also, as expected for thermal mechanisms (but also from melt or hydration), the peak in $R_{S,P}$ shifts to higher values in low velocity regions (median 1.9 for damping=35; Fig. S14). This trend is seen in both damped and undamped models.

4.2. Temperature estimates

As, to first order, the anomalies appear to be thermal in nature, we convert the low-velocity anomalies in the two main clusters (see boxes in Fig. S7) to temperature anomalies, using the same method as in *Civiero et al.* [2015]. To account for the fact that the tomography underresolves anomaly amplitudes and the recovered amplitudes depend on regularisation, we first scale (Fig. 7a) the S velocities according to the amplitude resolution with depth estimated from an idealised plume model (Fig. 3, Fig S15). The effect, and uncertainties of the amplitude resolution are discussed in detail in *Civiero et al.* [2015], where the results from differently damped inversions were in a similar range. Next, the scaled dV_S anomalies are converted to dT using a smooth

dV_S/dT derivative for a pyrolitic composition along a 1300°C adiabat (taken from *Styles et al.* [2011], Fig. S12). Uncertainties in this conversion to temperature anomalies are a few tens of degrees including consideration of uncertainties in the reference profile, and would largely work in the same direction for the conversion from P and S velocity anomalies, as both depend on shear modulus and density. Figure 7b shows that the inferred estimates of excess temperatures for both wave speeds agree well within their uncertainties and scatter, and fall between 100 and 200 °C.

Note that although the temperature estimates from dV_P and dV_S for both the Afar and MER clusters agree, the significant uncertainties and scatter would allow for some influence of other mechanisms. dT_S estimates tend to be slightly higher than dT_P above 500 km depth, while at larger depths dT_S tends to be somewhat lower than dT_P . The latter is probably due to the stronger decrease of the resolution with depth of V_S than V_P . The former may be a hint of a signature of deep melts. *Thompson et al.* [2015] used receiver functions to image a distinct low velocity zone above the transition zone beneath Afar. They interpreted this to be a melt layer caused by the release of volatiles from an upwelling below Afar. *Rooney et al.* [2012] also proposed a contribution of deep CO₂-assisted melting to the very slow velocities below Afar. Our tomographic model does not have the resolution to distinguish such local features [*Civiero et al.*, 2015], but does leave room for such an interpretation.

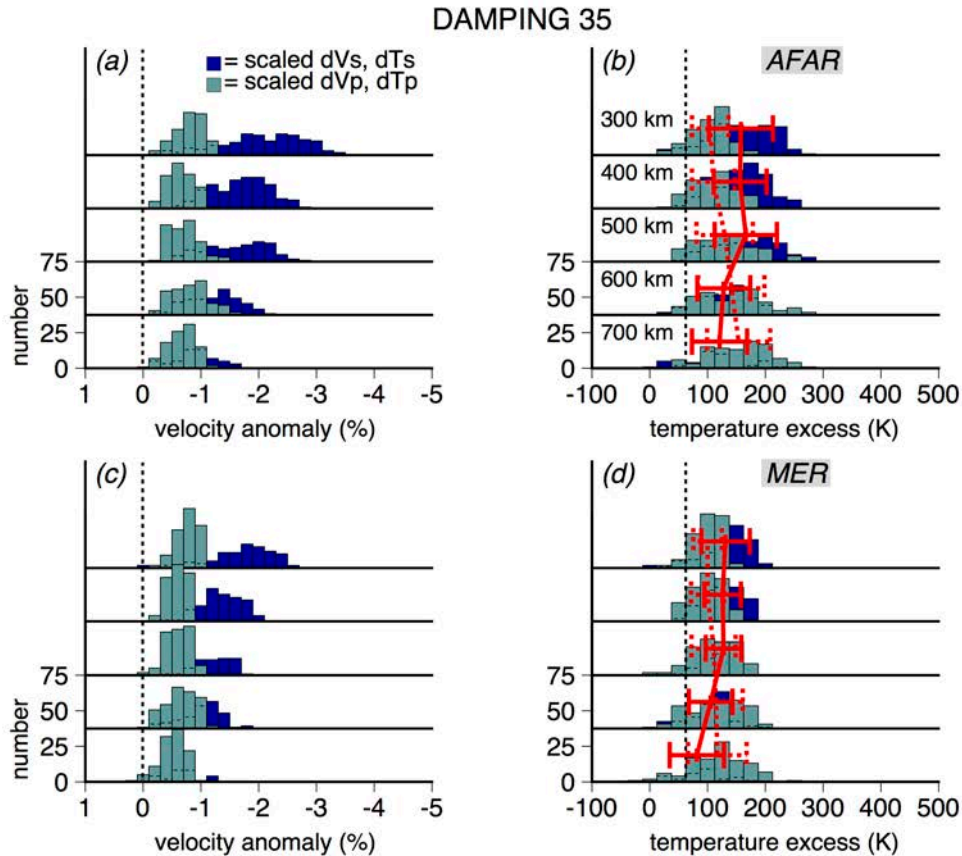


Figure 7. Histograms showing the range of velocity anomalies, after scaling for resolution, and the corresponding temperature anomalies from the NEAR-P15 P - and NEAR-S16 S /SKS-wave models (with a degree of damping of 35), at depths ranging from 300 to 700 km. The two regions considered are outlined by the boxes in Fig. S5 for P and Fig. S7 for S : Afar

Depression (a-b) and west of the MER (c-d). Panels a and c show in light green the P -wave velocity anomalies scaled according to the resolution as a function of depth based on the P -wave synthetic test in Civiero et al. [2015]; and in dark blue the S -SKS-wave velocity anomalies scaled according to the resolution as a function of depth based on the resolution test in Fig. 3, S15. Red dashed bars on panels b and d illustrate the mean and standard deviation of the temperature excess for P -velocity anomalies estimated in each region; red solid bars the mean and standard deviation of the temperature excess for S -velocity anomalies. These values show that temperatures inferred from dV_P and dV_S are in a similar thermal range of 100-200°C from 300 to 700 km depth.

Our temperature estimates are consistent with those inferred from geochemistry and petrology [Ferguson et al., 2013; Rooney et al., 2012], from seismic studies of the shallow mantle [Armitage et al., 2015; Rychert et al., 2012] and transition zone [Thompson et al., 2015], and from estimates of melt production [Armitage et al., 2015]. The estimates are lower than inferred below some other hotspots, e.g., Hawaii, North Atlantic and Deccan [Armitage et al., 2010; Watson and McKenzie, 1991]. They are also somewhat lower than those from Mulibo and Nyblade [2013], who in a recent tomographic study conclude that a thermal perturbation of ~150-300K reconciles P - and S -wave velocities beneath Tanzania and northwestern Zambia. The temperature estimates below the Northern EAR are consistent with the interpretation that the low velocities through the transition zone are due to warm thermal upwellings.

5. Conclusions

In this study, we combine S - and SKS-wave relative travel times to achieve a high-resolution tomographic shear-wave model, NEAR-S16, that extends throughout the whole upper mantle below of the Northern East-Africa. This S model complements a recent P travel time model, NEAR-P15, for the same region [Civiero et al., 2015]. The patterns of P - and S -wave velocities strongly resemble each other in terms of geometry and scale. The tomographic images reveal widespread low velocities in the upper 200 km of the mantle below the region. Below this depth, two clusters of low velocities emerge, situated below Afar and west of the MER. These extend through the transition zone. Both the P and S low velocity anomalies yield excess temperature estimates in the range of 100-200°C in these two clusters. These results give us more confidence in interpreting the structure imaged in both P - and S -SKS-wave models below the Northern East-African rift as being mainly the signature of multiple small-scale upper-mantle thermal upwellings.

A comparison of our imaged structure with that of previous global tomographic studies, reveals the consistency of our structures with these large-scale models, as shown in Fig. 8 for the global anisotropic S velocity model SGLOBE-RANI [Chang et al., 2015]. As more data have been added to the global models, the complexity of the low-velocity structure below the East African Rift has increased, e.g., compare S20RTS [Ritsema and van Heijst, 2000] which imaged a single large-scale feature that appeared to be continuous from the lower mantle below southern Africa to the upper mantle from Kenya/Tanzania to Afar, with S40RTS [Ritsema et al., 2011] and SGLOBE-RANI [Chang et al., 2015], which find two low-velocity anomalies in the upper mantle below Kenya/Tanzania and Afar, as well as variability within the lower-mantle low-velocity anomalies (Fig. 8). The comparison between our travel-time model and the global model SRANI clearly shows that what appeared like one large-

scale feature below Afar in several previous studies [e.g., Hansen *et al.*, 2012], has in fact internal small-scale structure, as was previously suggested based on surface wave tomography by Debayle *et al.* [2001]. Models S40RTS and SGLOBE-RANI even hint at a low-velocity anomaly at the base of the transition zone directly below the strong low velocities we image. As discussed fully in Civiero *et al.* [2015], this might be a ponding of warm material that forms the source of the small-scale upwellings we infer, a feature that is seen in analogue and numerical models of plume transition-zone interaction [e.g., Kumagai *et al.* [2007]; Tosi and Yuen [2011]]. Evaluation of the global models along different cross sections indicates that this low velocity region at the base of the transition zone may connect with one of the low-velocity features that appear to rise off the LLSVP below Africa. Further work to improve resolution of lower-mantle structure below the rift system, with more stations and/or by making use of a larger part of the waveforms will be necessary to identify the lower mantle source of the small-scale shallow mantle upwellings.

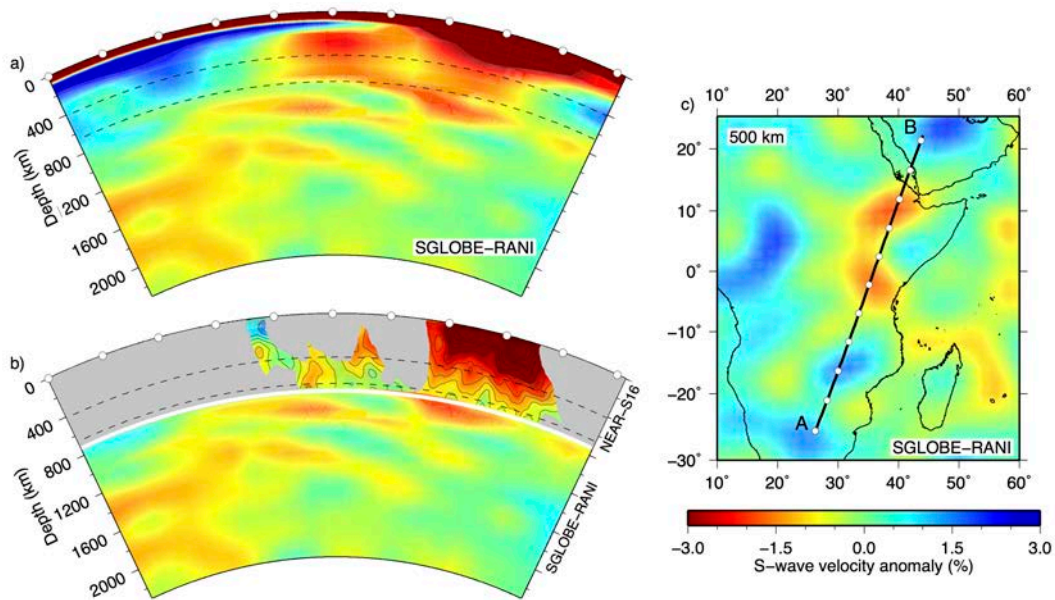


Figure 8. Comparison of global tomographic model SGLOBE-RANI [Chang *et al.*, 2015] and NEAR-S16. Map in (c) shows the location of the vertical cross section on a horizontal slice at 500 km depth through model SGLOBE-RANI. (a) vertical section through SGLOBE-RANI. White dots are spaced 10 degrees. (b) the top 800 km, our model NEAR-S16, structure below 800 km (marked by a white line) SGLOBE-RANI. NEAR-S16 is shaded gray where there are less than 5 rays per node. Note the consistency between the two models, with NEAR-S16 resolving more small-scale structure inside the large low velocity region in SGLOBE-RANI.

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Addis Ababa University (Contact for data: atalay.ayeale@aau.edu.et) and financial support for most of the Ethiopian permanent seismic stations comes from the International Science Program (ISP) of Uppsala University (Sweden). Other data was downloaded from the IRIS and GEOFON data centres. Funding for data collection was provided by NERC grants NE/E007414/1, and NE/J012297/1 and BHP-Billiton. We thank Ana Fereira for providing model SGLOBE-RANI and Ian Bastow and John Armitage for discussions. CC was supported by a Janet Watson Fellowship from the Department of Earth Science and Engineering at Imperial College, JH by NERC Fellowship NE/I020342/1, DK by NERC grant NE/L013932/1.

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Small-scale thermal upwellings under the Northern East African Rift from S travel-time tomography

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Introduction

The supporting information contains additional figures to further illustrate the resulting models and their resolution. In addition, it contains, to facilitate comparison, figures from our previously published P model, NEAR-P15 [Civiero *et al.*, 2015]. The supplementary table contains the details of all the stations and networks from which data were used in this study.

Part 1 – Surface wave model used as reference in damped inversions

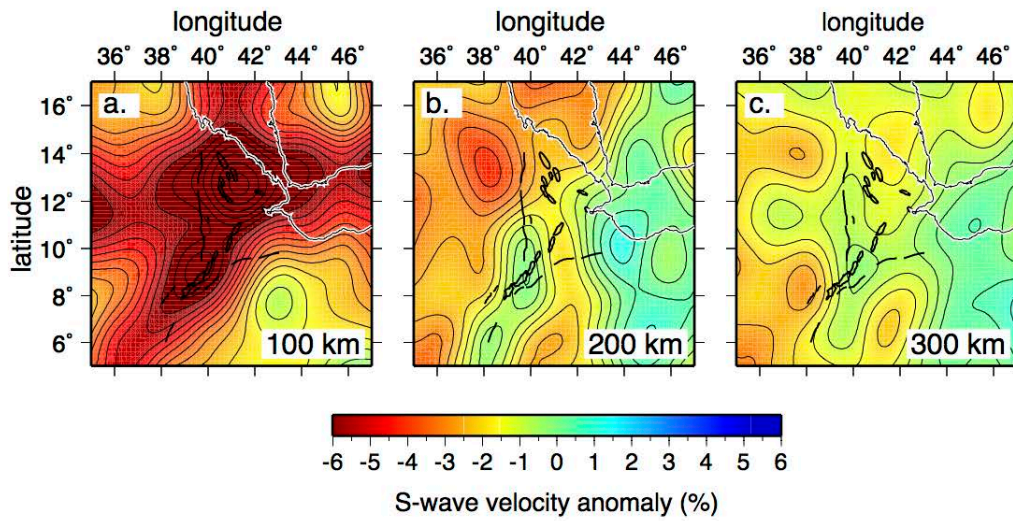


Figure S1. Depth slices showing the V_s anomalies from surface wave model from *Fishwick* (2010), which is used as 3-D reference model in the top 350 km for our damped inversions. The model covers all of Africa, but is only shown for our study region. The anomalies are relative to the average velocity-depth profile for the whole *Fishwick* (2010) model. The spacing between the contours is 0.50%. Note the anomaly scale is twice that of Fig S2 and others of our inverted model to allow displaying the structures at 100 km depth. Black lines delineate the major border faults and magmatic zones bounding the Afar Depression and black over white lines show coastlines.

Part 2 - Additional S inversions and P model

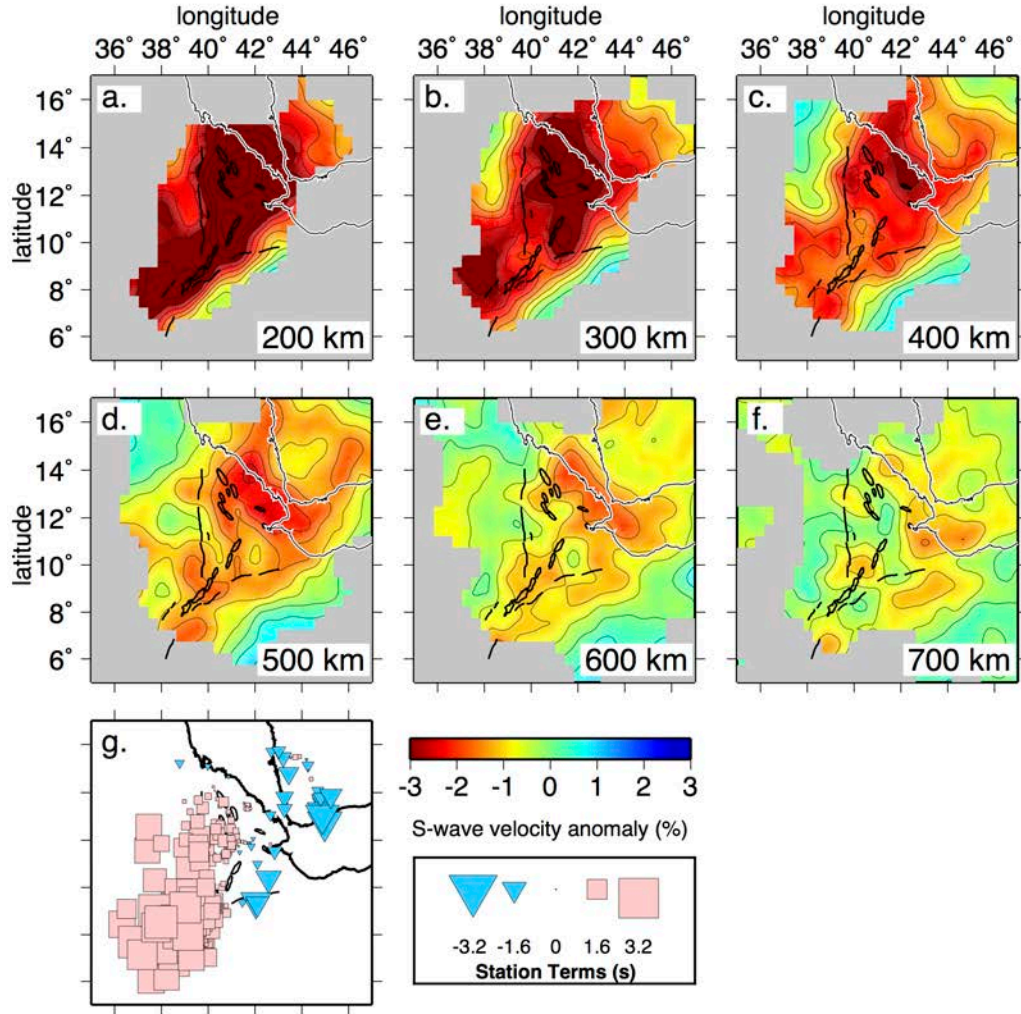


Figure S2. Depth slices through the *S-SKS-time preferred intermediate damped* tomographic model (flattening=4800, smoothing=153600, damping=35) at depths between 200 and 700 km. Regions with *less than 10 rays per node* are shaded gray. The spacing between the contours is 0.50%. Black lines delineate the major border faults and magmatic zones bounding the Afar Depression and black over white lines show coastlines. Triangular and square symbols in panel h represent the sign and magnitude of the station static terms. Comparison with Fig. 4 shows that the key features discussed in the text are all crossed by at least 10 rays per node.

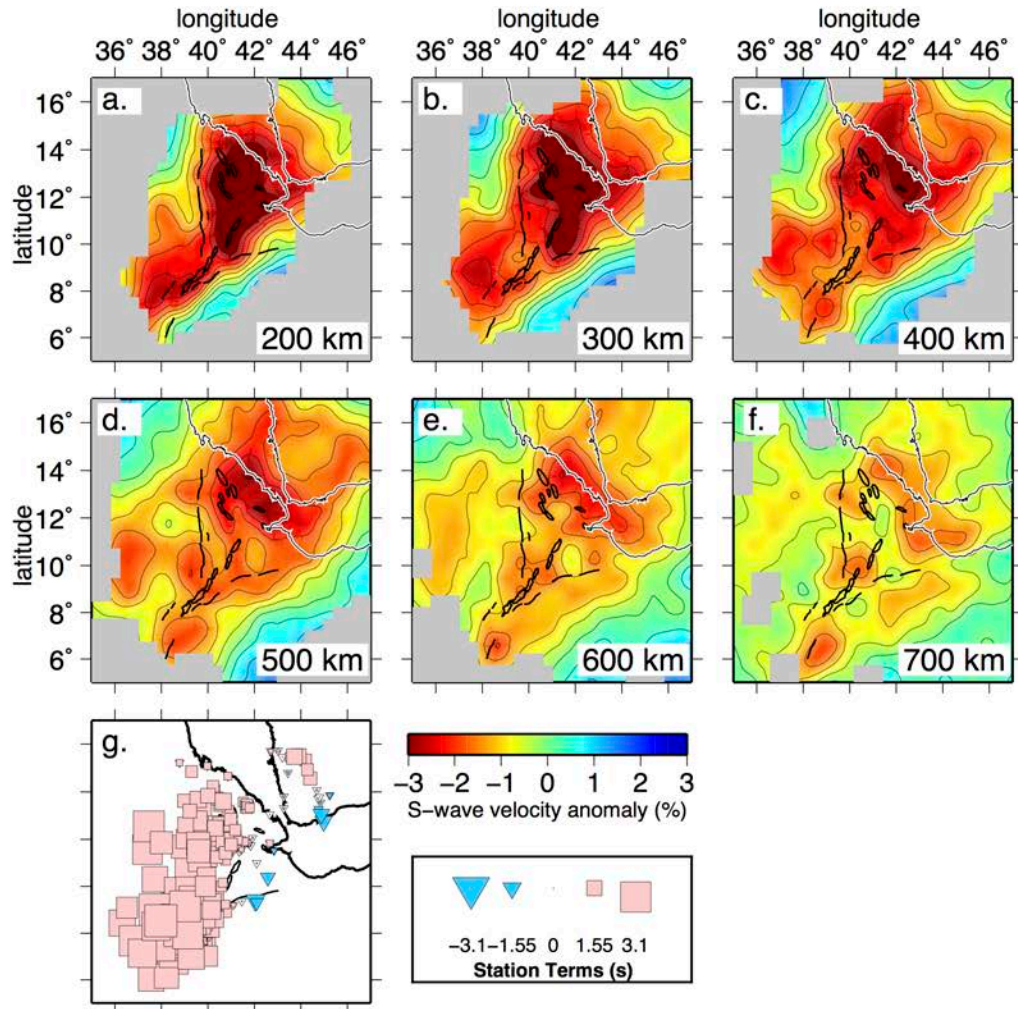


Figure S3. Depth slices through the *S-SKS-time undamped* tomographic model (flattening=4800, smoothing=153600, damping=0) at depths between 200 and 700 km. Regions with less than 5 rays per node are shaded gray. The spacing between the contours is 0.50%. Black lines delineate the major border faults and magmatic zones bounding the Afar Depression and black over white lines show coastlines. Triangular and square symbols in panel h represent the sign and magnitude of the station static terms. These depth slices illustrate that as in our preferred damping model in Fig. 4, the clusters of low-velocity anomalies that first appear around 200 km depth persist throughout the transition zone.

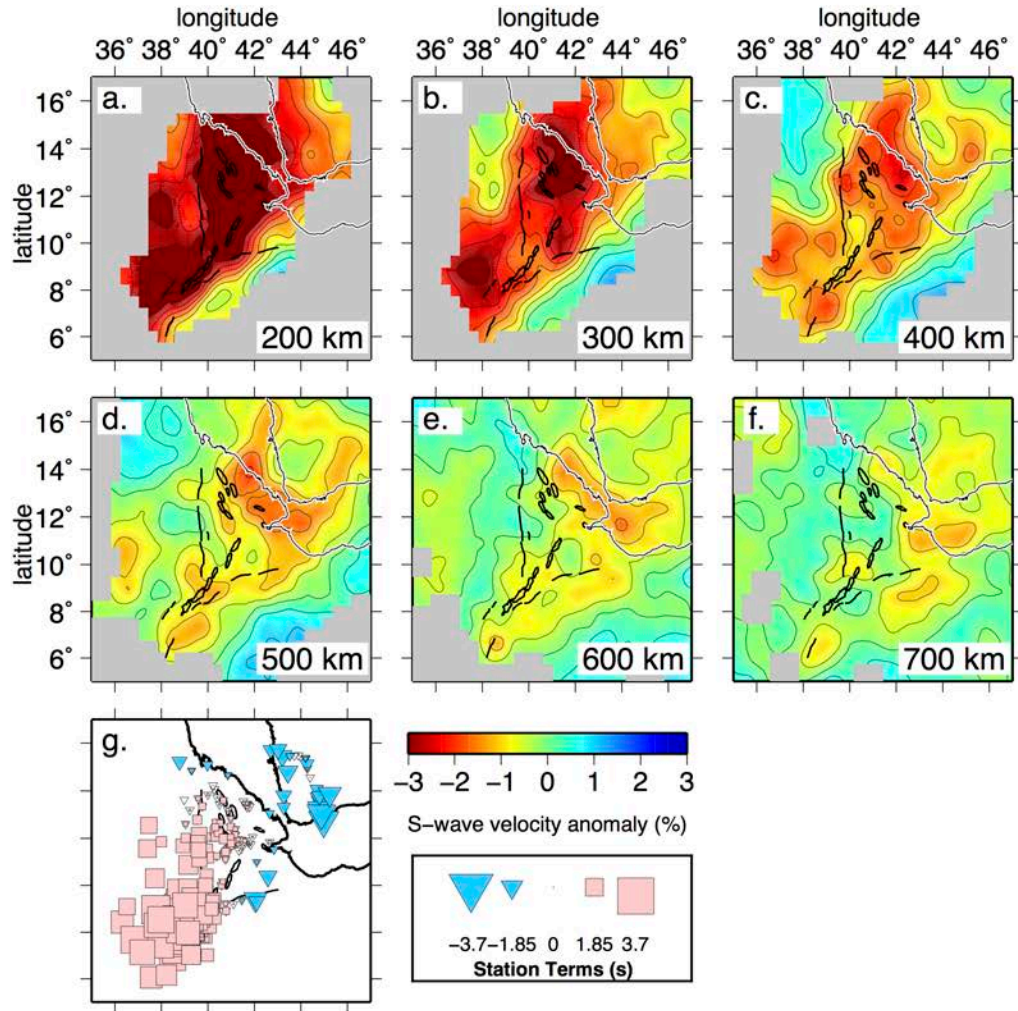


Figure S4. Depth slices through an S-SKS inversion with our preferred flattening (4800) and smoothing (153600) parameters but with a *stronger degree of damping* = 70. Triangular and square symbols in panel h represent the sign and magnitude of the station static terms. Regions with less than 5 rays per node are shaded grey. The spacing between the contours is 0.50%. Black lines delineate the major border faults and magmatic zones bounding the Afar Depression and black over white lines show coastlines. Structures are similar to those in our model with preferred damping (Fig. 4), but anomaly amplitudes decrease more strongly with depth.

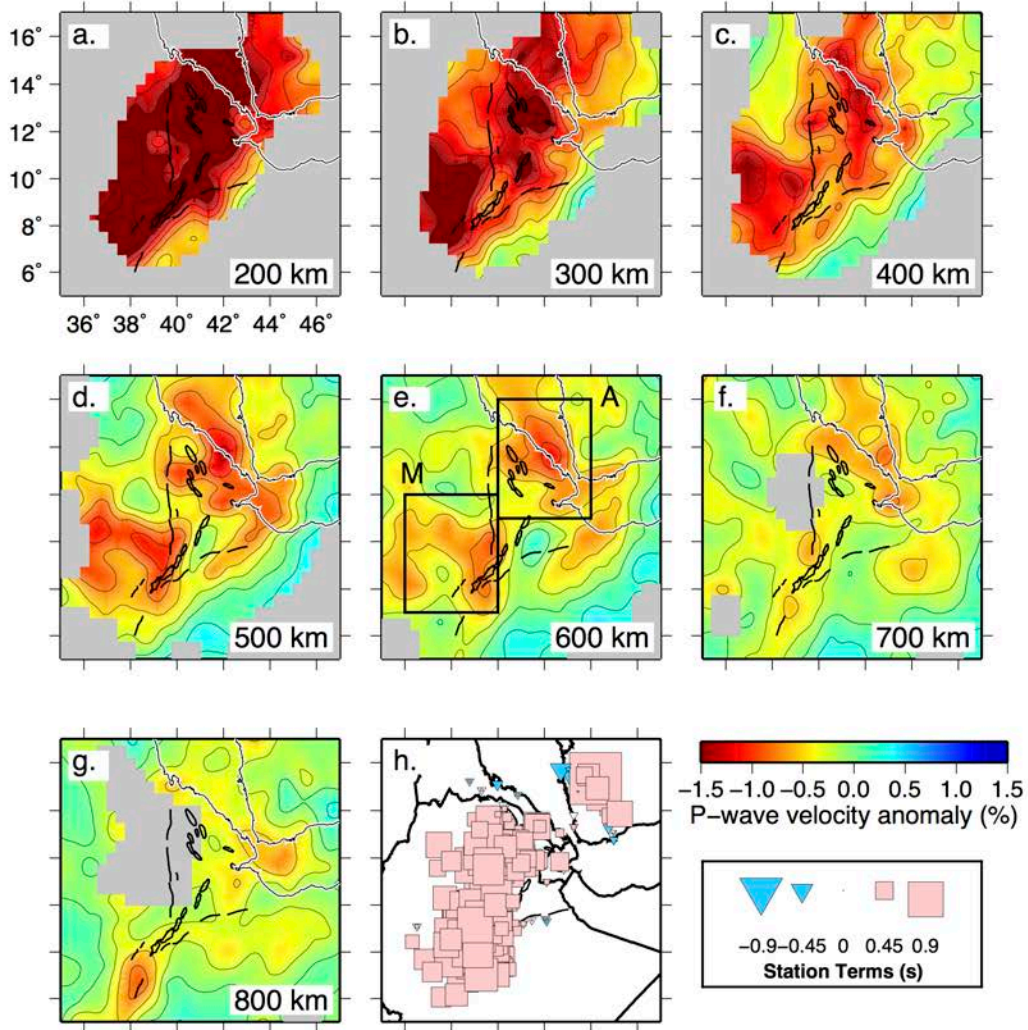


Figure S5. Depth slices through tomographic P model NEAR-P15 with preferred damping (flattening=4800, smoothing=153600, damping = 35) from *Civiero et al.* [2015], at depths between 200 and 800 km. Triangular and square symbols in panel h represent the sign and magnitude of the station static terms. Regions with less than 5 rays per node are shaded grey. The spacing between the contours is 0.25%. Black lines delineate the major border faults and magmatic zones bounding the Afar Depression and black over white lines show coastlines. The two boxes covering Afar (A) and an area west of the MER (M) in panel e. show the regions used in our temperature interpretation in Fig. 7. The structures in this model are very similar to those in NEAR-S16 (Fig. 4).

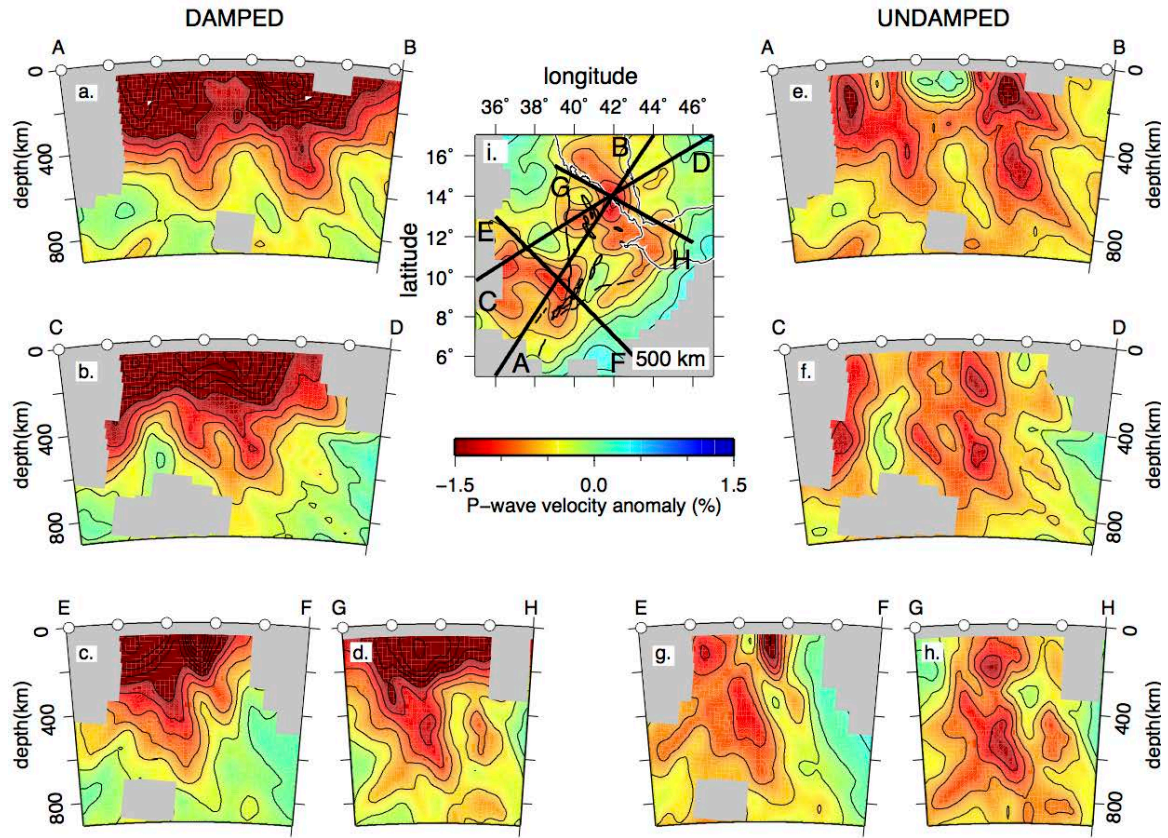


Figure S6. Vertical cross sections through the tomographic *P* model NEAR-P15 from Civiero *et al.* [2015] for preferred (damping=35) (a-d) and undamped (damping=0) cases (e-h) (both flattening=4800, smoothing=153600). The location of the cross-sections (black lines) is shown in the 500 km depth slice through the damped model (i). Regions with less than 5 rays per node are shaded grey. The spacing between the contours is 0.25%. White points indicate the distance every 2 degrees. Cross section A-B cuts through subvertical downwellings below Afar and west of the MER. C-D is a cross section through the prominent low-velocity anomalies in the Afar region and directly next to the MER. Section E-F cross cuts the anomaly next to the MER, while section G-H provides another view of the Afar low-velocity anomaly. The undamped models (e-h) illustrate the 100-200 km width of the low-velocity structures, while the damped models (a-d) emphasize the continuity between shallow and transition zone structure. Comparison with the *S* models (Fig. 5) reveals the similarity of the *P* and *S* structures.

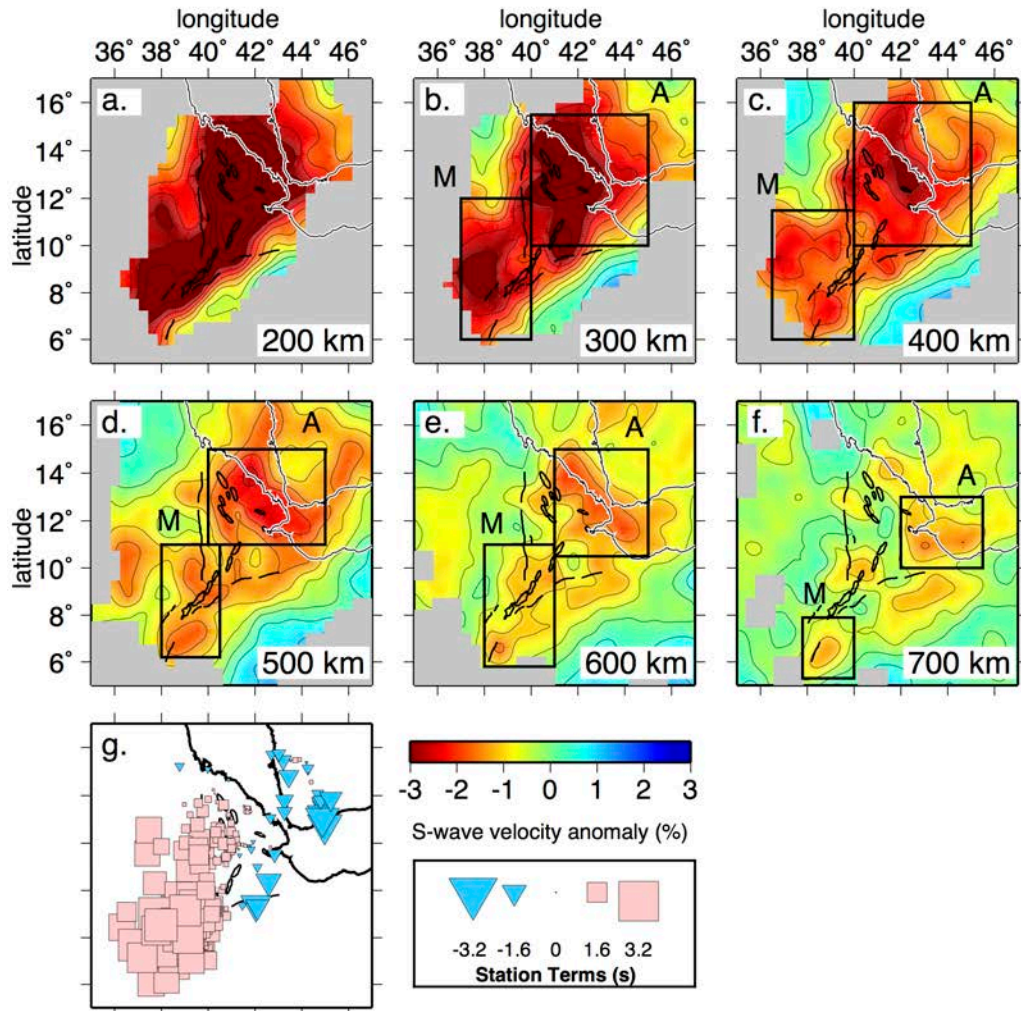


Figure S7. Depth slices through tomographic *S* model NEAR-S16 with preferred damping (flattening=4800, smoothing=153600, damping = 35). Same as Fig. 4, but with superimposed a set of boxes covering Afar (A) and an area west of the MER (M) that delineate the regions used in our temperature interpretation in Fig. 7.

Part 3 – Additional resolution tests

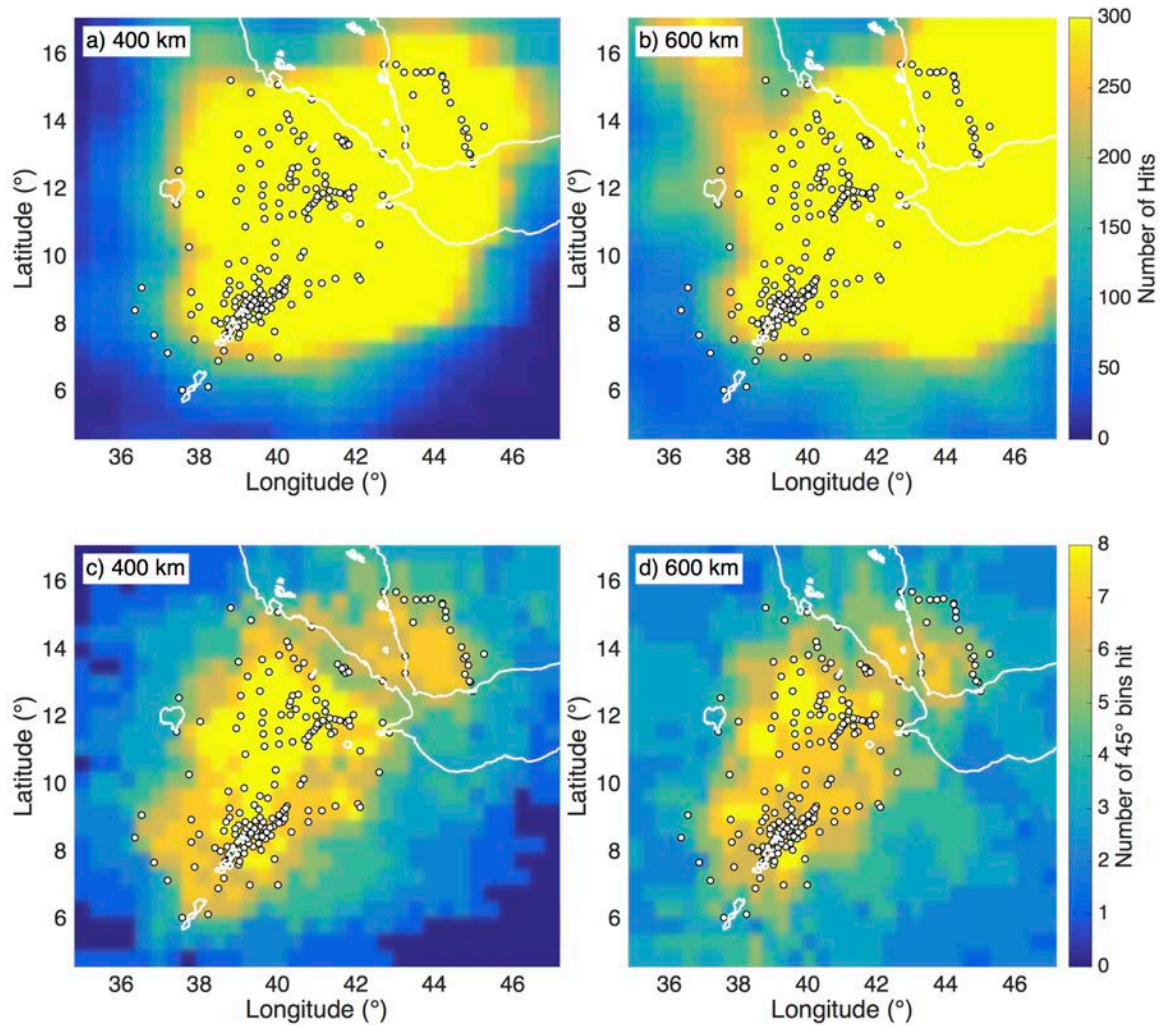


Figure S8. Maps comparing hit count (top) and crossing rays (bottom) at a depth of 400 km (a,c) , and 600 km (b,d) for the north-east East-African Rift. The azimuthal coverage is expressed as the number of different 45° bins that are crossed, where 8 would be a complete 360° coverage. a,b) Number of hits per node based on the rays used in this study at 400 and 600 km depth, respectively. c,d) Number of different 45° bins hit by rays at 400 and 600 km depth, respectively. Note the good coverage both in number of hits and crossing of rays. Coast lines and lakes are shown in white. White circles show seismic stations used in each study.

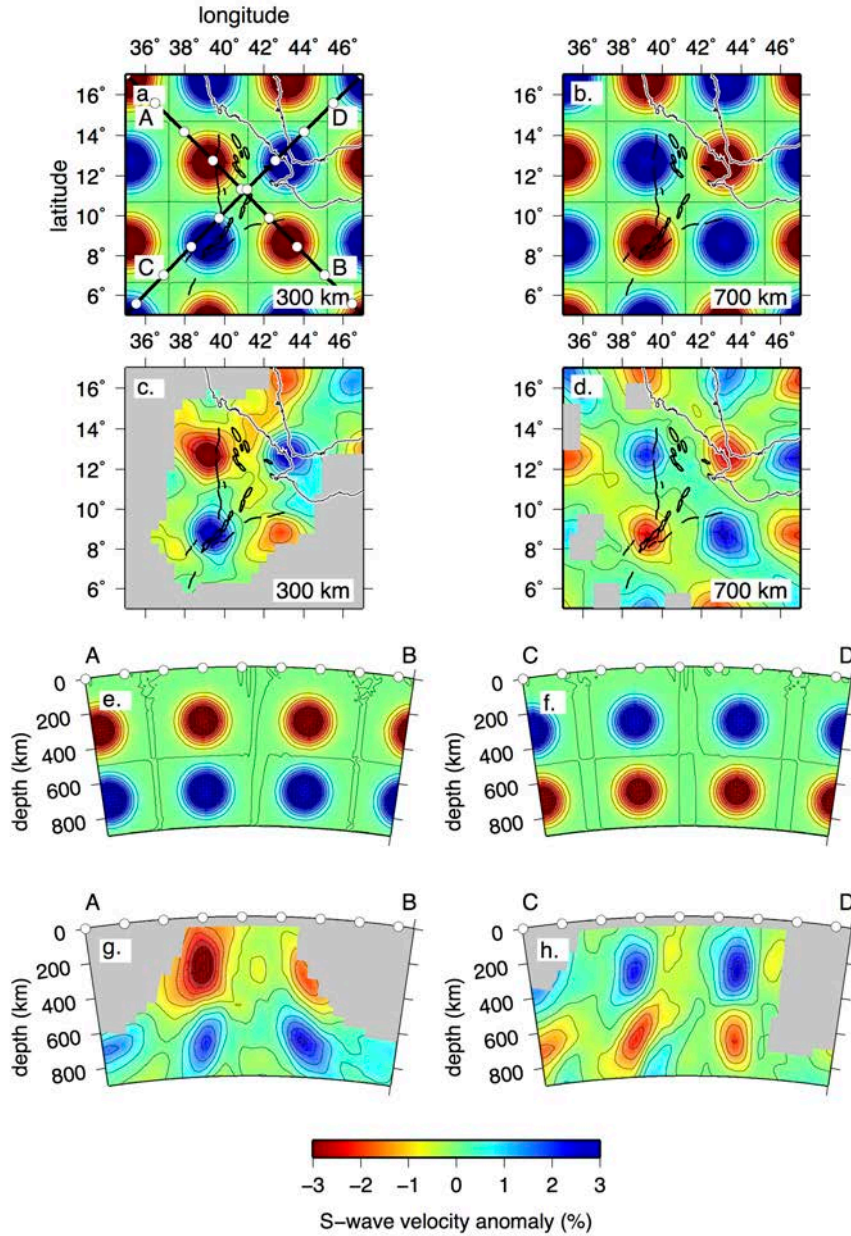


Figure S9. Checkerboard resolution tests for the S-SKS-times tomographic inversion, using 250 km wide spherical anomalies with a Gaussian amplitude profile, peaking at 7% (width defined as the distance to 20% of the maximum amplitude). The same raypaths and inversion parameters as in the data inversion are used and Gaussian noise of 0.37 s is added to mimic that in the data. Damping is not included, as this does not affect the spatial distribution of resolution. Regions with less than 5 rays per node are shaded grey. Contour interval is 0.50%. White circles along the top of the vertical profiles are spaced every 2°. (a, b) input model at 300 and 700 km depth. (c, d) Recovered checkers at 300 and 700 km depth. (e-h) vertical cross sections, oriented approximately rift-perpendicular (e,g) and approximately parallel to the rift trend (f,h), through the input (e,f), and output (g,h) models (orientations of the profiles are shown in depth slice a).

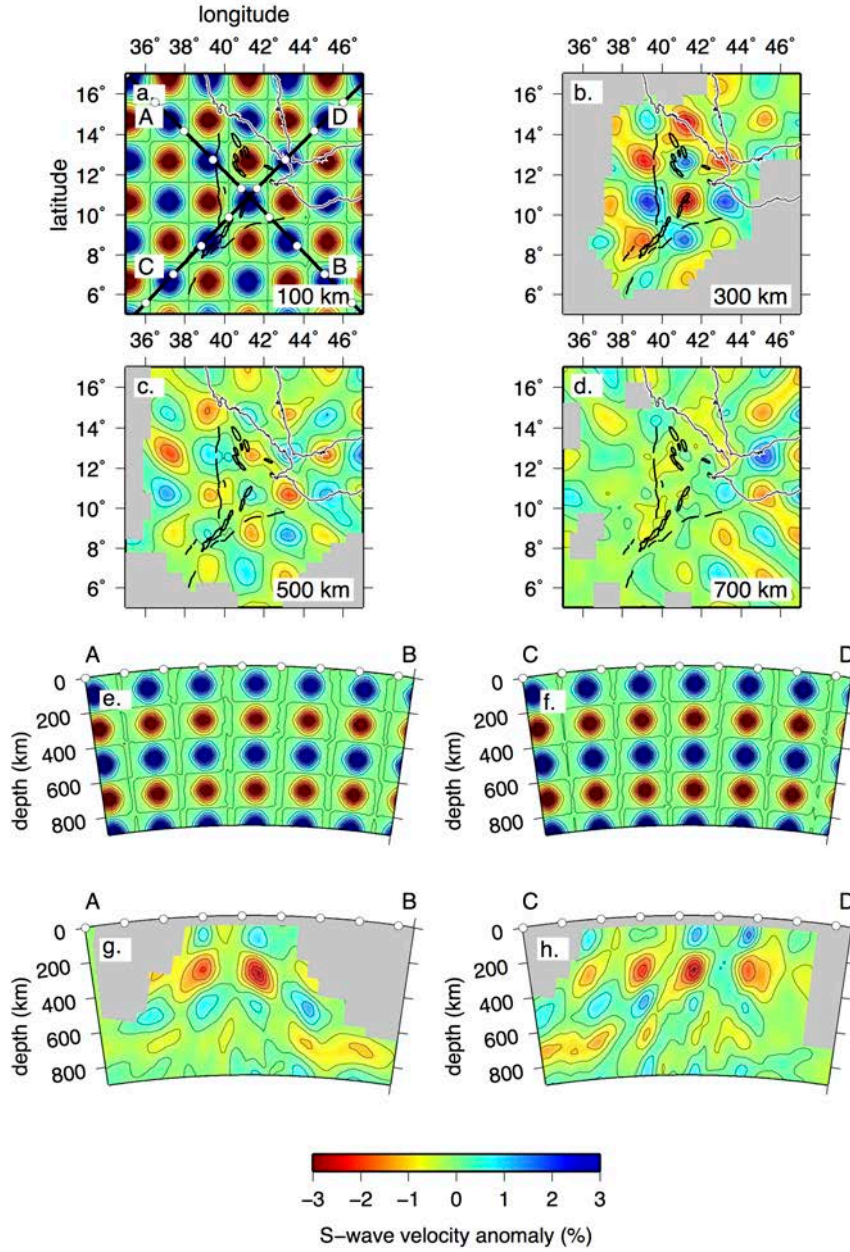


Figure S10. Checkerboard resolution test for the S-SKS inversion, as in Fig. S9, but with checkers of 125 km width, again with 7% peak amplitudes. (a) example of the input model at 300 km depth. (b,c,d) Recovered checkers at 300, 500 and 700 km depth, respectively. Panels (e-h) show vertical cross sections through the checker structure, oriented approximately rift-perpendicular (e,g) and rift-parallel (f,h), through the input (e,f), and output (g,h) models (orientations of the profiles are shown in depth slice a).

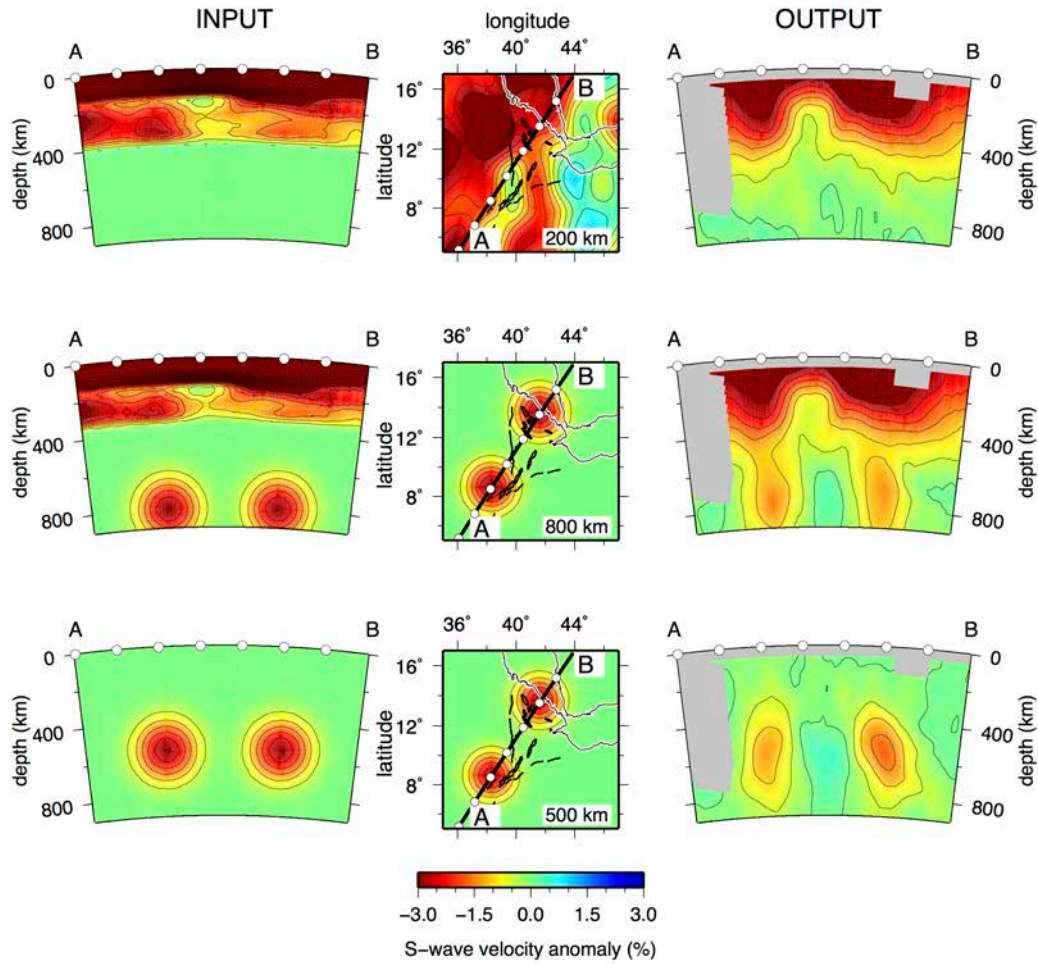


Figure S11. Three resolution tests along cross-section A-B for the S-SKS inversion; orientation is shown in the depth slices through each input model in the middle column (note the different depths of the slices for the different models). Regions with less than 5 rays per node are shaded gray. The spacing between the contours is 0.50%. White circles along the top of the vertical profiles mark the distance every 2°. The left-hand column shows synthetic input models and the right-hand column the corresponding recovered output models. The synthetic inversion is done using the same parameters as for the data-derived inversion. The synthetic test in the first row uses the S-structure estimated from the surface-wave model from Fishwick [2010] down to 350 km depth as the input model. The panels in the second row show a test with the same surface-wave-derived structure at shallow depths, plus a set of Gaussian low-velocity anomalies with peak amplitudes of 3.2% and a 380 km diameter (defined as the distance to 20% of the maximum amplitude) along line A-B, placed beneath the transition zone (centres at 800 km depth), to represent lower-mantle structure with a large excess temperature ($\sim 400^\circ\text{C}$). The third row shows a test using two low-velocity anomalies with 3.2% maximum-amplitude and 380 km diameter positioned within the transition zone, centred at 550 km depth. While the shallow and transition-zone structures are quite well recovered, the lower mantle structures are not as well resolved and smeared upwards.

Part 4– Seismic ratios and thermal interpretation

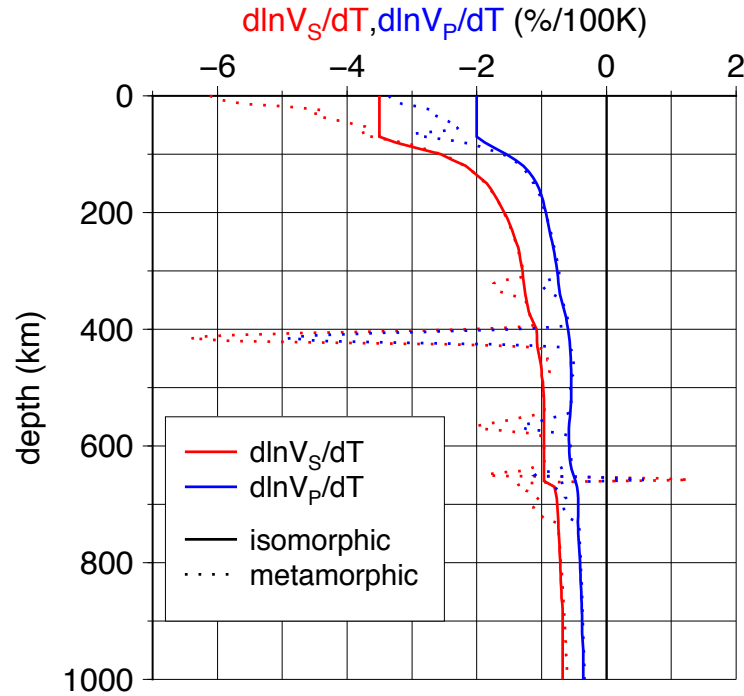


Figure S12. P and S velocity-temperature derivatives used to set up our synthetic plume resolution tests (Fig. 3) and perform our conversion of imaged velocity to temperature anomalies (Fig. 7). The dotted profiles are full (metamorphic) derivatives that include the effects of phase transitions. They were computed using PerPleX [Connolly, 2005] along a 1300°C adiabat for a pyrolite composition using mineral parameters from database stx08 [Xu *et al.*, 2008], with composite attenuation model Qg (above 400 km) [Van Wijk *et al.*, 2008] and Q6 (below) [Goes *et al.*, 2004]. The solid profiles correspond to $dV_{P,S}/dT$ derivatives that were smoothed to represent the isomorphic derivatives without the effects of phase boundary topography. For our scaling, we use the isomorphic derivatives because the differential travel-time tomography cannot resolve localized phase boundary anomalies [Civiero *et al.*, 2015]. Above 70 km depth, the isomorphic scaling is set to a constant value of -2% per 100 K for V_P and 4% per 100 K for V_S , more representative for a (cooler) lithospheric geotherm.

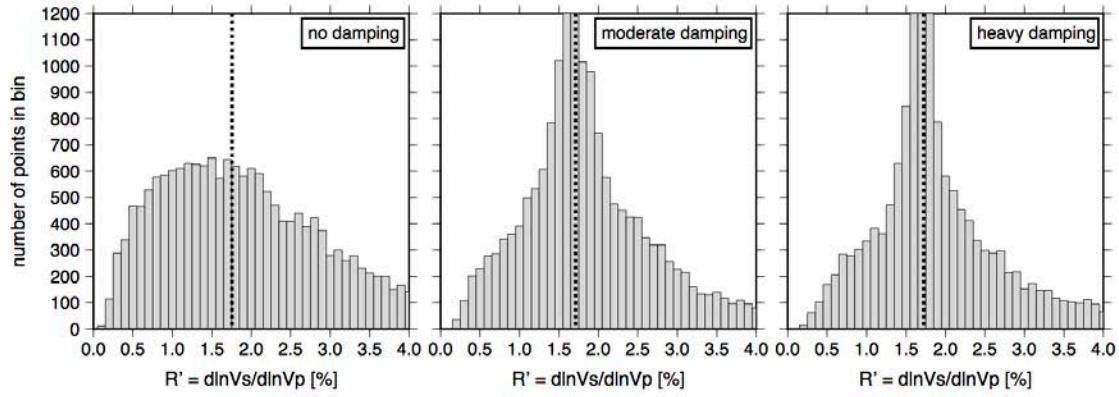


Figure S13. $R_{S,P}$ ($=\ln V_S/\ln V_P$) distribution estimated from the NEAR-P15 P - and NEAR-S16 S -velocity models for the whole region of interest, from inversions with our preferred regularisation (flattening=4800, smoothing=153600 for both V_P and V_S) and different degrees of damping (0-35-70). For all the three differently damped models, the $R_{S,P}$ distribution is peaked around ~ 1.7 (median value), consistent with a dominantly thermal origin of the anomalies. As the S to P conversion used for the 3-D starting velocity model assumed the anomalies were thermal, increasing the damping parameter the $R_{S,P}$ distribution leads to a distribution more strongly peaked around this value. For each degree of damping, a large quantity of scatter around is also present and may reflect non-thermal effects and/or resolution differences between the P and S models.

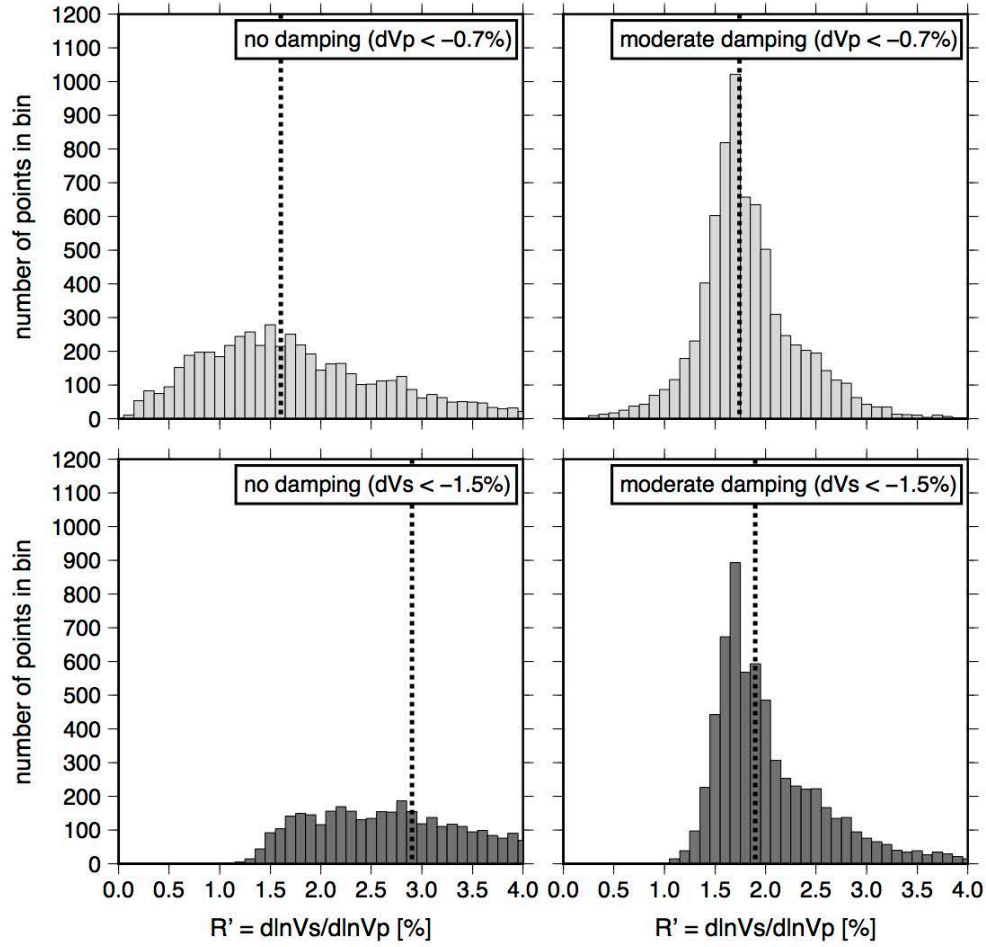


Figure S14. $R_{S,P}$ distributions calculated only for the lowest NEAR-P15 P -wave velocities (top panels) and NEAR-S16 S -wave velocities (bottom panels) ($dV_p < -0.7\%$ and $dV_s < -1.5\%$ respectively) in undamped models (first column) and moderately damped (damping=35) models (second column). $R_{S,P}$ distributions shift towards higher values compared to the distributions that include slow and fast anomalies in Fig. S13, as might be expected if the anomalies are thermal in nature, but also if some of the low velocities are due to the presence of fluids/melt.

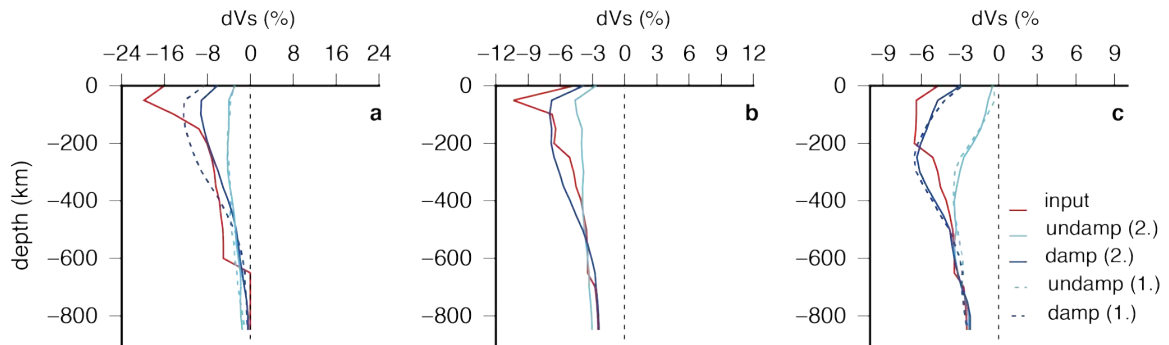


Figure S15. Profiles through the centre of the three synthetic plume models from Fig.3, a) Superplume, b) single plume, c) double plumes, showing the input and retrieved anomalies. Locations for plumes 1 (approximately MER) and 2 (Afar) for case c are shown on panel m of Fig. 3.

Part 5 – Station information

<i>Stat</i>	<i>Network</i>	<i>Lat(°)</i>	<i>Long(°)</i>	<i>Elev(km)</i>
AAIR	YI	-9.95	33.9	0.53
AAUS	AF,XI	9.03	38.77	2.25
ABAE	ZF,ZR	13.35	39.76	1.45
AD02	ZK	11.35	40.69	0.52
AD04	ZK	11.54	40.84	0.48
AD05	ZK	11.61	40.91	0.45
AD07	ZK	11.73	40.99	0.40
AD08	ZK	11.78	41.03	0.39
AD09	ZK	11.82	41.05	0.39
AD10	ZK	11.90	41.14	0.37
AD11	ZK	11.74	41.3	0.37
AD14	ZK	11.94	41.45	0.42
AD15	ZK	11.88	41.71	0.09
AD16	ZK	11.82	41.75	0.12
AD17	ZK	11.74	41.84	0.15
AD18	ZK	11.91	41.79	0.45
ADBA	XW	13.55	44.84	0.70
ADEE	YJ	7.79	39.91	2.48
ADEN	XW	12.78	44.98	0.06
ADHO	YR	17.24	54.28	0.91
ADTE	ZF	11.12	40.76	0.51
ADUA	XW	15.00	48.97	1.38
ADUE	YJ	8.54	38.9	1.75
ADYE	ZF	13.64	38.98	1.86
AFME	ZE,ZR	13.20	40.86	-0.06
AHME	ZR	14.09	40.28	0.05
AKEE	ZF	10.89	39.17	3.23
ALE	YR	9.42	42.03	2.03
ALGU	XW	13.05	44.93	0.1
AMBA	XW	-8.11	33.26	1.42
AMME	YJ	8.30	39.09	1.67
ANGA	XI	-2.50	36.8	0.00
ANID	XW	15.47	43.2	0.15
ANKE	YJ	9.59	39.73	2.98
ARBA	XI	6.07	37.56	1.27
ARCH	YI	- 10.02	33.93	0.5
AREE	YJ	8.94	39.42	1.83

ARUT	XW	15.16	51.03	0.02
ASE	7C	11.00	42.1	0.36
ASSE	YR	13.06	42.65	0.02
ASYE	ZF,ZR	11.56	41.44	0.37
ATD	G	11.53	42.85	0.61
AWRE	YZ	12.07	40.07	0.85
AYDO	YR	16.99	53.36	0.04
AYNO	YR	17.26	53.89	0.86
BAHI	XI	11.57	37.39	1.79
BANO	YR	17.69	54.44	0.46
BARE	ZE	12.64	40.36	0.34
BARI	XI	0.47	35.98	1.01
BEDE	YJ	8.91	40.77	1.71
BELA	XI	6.93	38.47	1.92
BEND	ZP	0.58	31.39	1.35
BERE	ZE	12.17	41.19	0.57
BIHA	ZP	-2.64	31.32	1.46
BIRH	XI	9.67	39.53	2.81
BKBA	ZP	-1.36	31.81	1.27
BOBE	ZF	10.38	42.57	0.94
BOKO	XI	-2.26	37.73	0.98
BORE	YJ	8.75	39.55	1.25
BOVE	YZ	12.66	40.52	0.76
BREE	ZE	12.17	41.19	0.58
BTIE	ZF,ZR	11.19	40.02	1.66
BURO	UN	0.86	30.17	0.98
BUTE	YJ	8.12	38.38	2.09
BUTI	ZP	1.82	31.33	0.62
C01	Z5	1.36	29.76	0.95
C02	Z5	0.85	29.61	1.08
C03	Z5	0.7	29.52	1.05
C04	Z5	0.66	29.88	0.85
C05	Z5	0.48	29.5	1.20
C06	Z5	0.32	29.75	1.18
C07	Z5	0.29	29.34	1.18
C08	Z5	0.12	29.28	1.69
C09	Z5	-0.14	29.6	0.95
C10	Z5	-0.16	29.23	1.87
CAYE	YR	14.86	39.31	2.44

CHAE	YJ	9.31	38.76	2.65
CHEF	XI	6.16	38.21	1.70
CHIE	ZE	11.6	40.02	0.94
CHIM	ZP	-8.83	34.03	1.10
CLIN	YI	-9.96	33.81	0.54
DABI	XW	15.13	44.27	2.38
DAHO	YR	17.53	54.35	0.54
DALA	XW	13.73	44.74	1.44
DALE	ZR	14.23	40.22	-0.10
DAME	ZR	11.69	40.96	0.42
DAMT	XW	14.09	44.68	1.90
DAMY	GE	14.57	44.39	2.49
DEBE	YJ	8.78	39	1.91
DELE	XI	8.44	36.33	1.97
DERU	XW	16.84	51.83	0.88
DICE	ZF	11.91	41.57	0.46
DIGE	ZE	12.33	40.27	0.68
DIKE	YJ	8.06	39.56	2.75
DKUM	XW	13.27	44.76	0.4
DMRK	XI	10.31	37.73	2.36
DMTO	YR	17.73	55.07	0.44
DODT	AF	-6.19	35.75	1.11
DOLE	YR	15.1	39.98	0.09
DONE	YJ	8.51	39.55	1.31
DSS	YR	11.12	39.64	2.55
E31	XJ	8.78	39.86	1.01
E33	XJ	8.93	39.93	0.98
E34	XJ	7.21	38.6	1.93
E35	XJ	9.13	40.17	0.85
E36	XJ	9.11	40.01	0.77
E37	XJ	8.17	38.7	1.8
E39	XJ	9.24	40.13	0.77
E40	XJ	9.36	40.22	0.74
E41	XJ	8.01	38.53	1.91
E42	XJ	8.88	40.1	1.06
E43	XJ	9.25	39.5	3.29
E46	XJ	8.71	39.69	1.24
E47	XJ	8.46	39.45	1.45
E48	XJ	7.62	38.99	2.61
E50	XJ	8.27	39.5	2.07
E51	XJ	8.15	39.35	2.08

E53	XJ	8.04	39.01	1.70
E54	XJ	8.12	39.14	2.08
E55	XJ	8.3	38.95	1.68
E56	XJ	8.46	39.06	1.64
E57	XJ	8.58	39.13	1.82
E58	XJ	8.69	39.18	2.06
E59	XJ	8.71	39.35	1.68
E60	XJ	8.62	39.45	1.63
E61	XJ	8.9	39.62	1.16
E63	XJ	8.26	39.24	1.78
E65	XJ	8.4	39.21	1.55
E66	XJ	9.03	39.53	1.72
E67	XJ	8.38	39.68	2.14
E68	XJ	8.78	39.26	2.29
E69	XJ	7.93	38.72	1.68
E70	XJ	8.88	39.15	2.23
E71	XJ	8.69	38.9	1.98
E72	XJ	8.49	39.83	1.58
E73	XJ	7.74	39.03	2.5
E75	XJ	7.91	38.95	1.75
E76	XJ	7.72	38.65	1.67
E77	XJ	7.86	38.79	1.67
E78	XJ	8.59	39.7	1.22
E79	XJ	7.63	38.71	1.59
E80	XJ	8.48	39.31	1.66
E82	XJ	8.85	40.01	0.97
E83	XJ	7.8	38.79	1.90
E84	XJ	8.7	39.4	1.54
E85	XJ	8.46	39.59	1.32
EITE	YR	15.24	38.78	2.17
ELLE	ZF	11.26	40.38	0.67
ERTE	ZF,ZR	13.45	40.5	-0.01
EYUN	XW	14.78	49.27	0.18
FAME	YR	13.57	41.52	0.62
FASH	XW	15.44	50.95	0.14
FICH	XI	9.78	38.74	2.83
FINE	ZE,ZR	12.07	40.32	0.78
FOPO	ZP	0.66	30.28	1.53
FURI	IU	8.9	38.68	2.56
GALE	ZR	13.73	40.39	-0.09
GASE	ZF	11.68	38.92	2.97

GDR	YR	12.56	37.45	2.1
GEAN	XW	16.71	49.52	0.97
GEIT	ZP	-2.88	32.22	1.28
GEWE	YJ,ZF	10	40.57	0.6
GHAD	XW	16.25	52.21	0.05
GHDI	XW	15.64	52.16	0.03
GLUM	ZU	-2.62	36.19	1.30
GOBA	XI	7.03	39.98	2.73
GOMA	XD	-4.84	29.69	0.88
GTFE	YJ	9	39.84	1.04
GUDE	XI	8.97	37.77	2.02
GULE	ZR	13.69	39.59	2.02
HADO	YR	17.22	55.19	0.09
HAHY	XW	15.21	49.09	1.08
HALE	ZF,ZR	13.84	40.01	0.23
HAMA	ZP	-3.83	32.64	1.23
HATT	XW	17.32	52.11	0.77
HAYO	YR	17.18	53.34	0.83
HERO	XI	7.03	39.28	2.37
HIRN	XI	9.22	41.11	1.82
HOSA	XI	7.56	37.86	2.31
HOTA	XW	13.06	44.88	0.13
HUMY	UN	0.76	30.04	1.00
HYNE	ZF	9.31	42.1	1.98
IGRE	ZR	12.25	40.46	0.68
INEE	YJ	9.9	39.14	2.69
IRIN	ZP	-7.76	35.69	1.56
ITOJ	UN	0.84	30.23	1.00
JIMA	XI	7.68	36.83	1.66
JNJA	ZP	0.45	33.18	1.13
KABA	UN	0.78	30.13	0.92
KABE	UN	0.87	30.47	1.3
KAGO	UN	0.68	30.46	1.52
KAKA	XI	0.56	34.8	1.48
KARE	XW	17.16	51.93	0.09
KARE	YJ	10.42	39.93	0.86
KARU	UN	0.79	30.22	1.11
KAS2	UN	-0.03	30.15	0.92
KASS	UN	0.58	30.31	1.50
KBLE	ZP	-1.25	29.99	1.88
KERM	ZU	-2.83	35.98	1.14

KGMA	ZP	-4.88	29.63	0.82
KHAW	XW	13.81	43.25	0.01
KHLA	XW	13.8	44.81	1.46
KIBA	XD	-5.32	36.57	1.50
KIBE	XD	-5.38	37.48	1.00
KIBO	ZP	-3.58	30.71	1.49
KIG	AF	-1.96	30.06	1.54
KILE	UN	0.21	30.01	1.35
KINY	UN	0.51	30.13	1.70
KISA	UN	0.59	30.74	1.29
KITU	XI	-1.37	38.00	1.13
KMBO	GE	-1.13	37.25	1.94
KMTW	UN	0.74	30.38	1.56
KOBE	ZF	12.15	39.63	1.51
KOMO	XD	-3.84	36.72	1.11
KOND	XD	-4.90	35.8	1.42
KOTE	YJ	9.39	39.4	2.87
KOZE	ZR	12.49	40.98	0.54
KR42	XI	0.04	35.73	2.16
KTWE	AF	- 12.81	28.21	1.23
KYLA	ZP	-9.60	33.87	0.50
LAEL	ZP	-8.57	32.06	1.60
LALE	ZF	12.03	39.04	2.42
LAVE	ZR	13.60	40.66	0.60
LBB	AF	- 11.63	27.49	1.28
LEME	YJ	8.61	38.61	2.11
LODK	GE	3.42	35.36	0.67
LONG	XD	-2.73	36.7	1.38
LOSS	ZP	-8.42	33.16	1.20
LSZ	IU	- 15.28	28.19	1.2
LUGH	YI	- 10.03	32.83	0.59
LULE	ZR	11.99	40.7	0.59
LUSA	XW	16.49	52.57	0.05
LWND	ZF	-2.75	36.04	1.07
LYDE	ZF	12.05	41.93	0.43
MADO	YR	17.2	54.38	0.83
MAFI	ZP	-8.31	35.31	1.87
MAKA	ZP	-8.85	34.83	1.69
MALE	ZP	1.07	34.17	1.13

MAUS	ZP	-2.74	36.7	1.33
MAWI	XW	15.47	43.52	1.88
MAYE	ZF,ZR	12.78	39.53	2.44
MBAR	II	-0.6	30.74	1.39
MBWE	XD	-4.96	34.35	1.1
MDYO	YR	17.46	53.36	0.57
MECE	YJ	8.59	40.32	1.77
MEGE	ZE	11.49	41.34	0.35
MEKE	YJ	8.16	38.83	1.90
MELE	YJ	9.31	40.2	0.76
MGOR	ZP	-6.83	37.67	0.50
MIKU	ZP	-7.40	36.99	0.52
MILE	ZE	11.42	40.76	0.49
MIRA	UN	0.66	30.57	1.38
MISE	ZF	9.24	40.76	1.31
MITU	XD	-6.02	34.06	1.57
MKRE	ZP	-4.28	30.42	1.18
MLBA	ZP	-1.84	31.67	1.34
MOKA	XW	13.31	43.26	0.03
MSEY	II	-4.67	55.48	0.47
MTOR	XD	-5.25	35.4	1.10
MUGO	YR	16.90	53.77	0.04
MUKA	XW	14.49	49.04	0.04
MWEY	UN	-0.19	29.9	0.96
MZM	AF	-11.43	34.03	1.26
NAB1	YW	13.39	41.66	1.33
NAB2	YW	13.43	41.71	1.21
NAB3	YW	13.38	41.75	1.28
NAB4	YW	13.48	41.68	0.70
NAB5	YW	13.32	41.71	1.27
NAB6	YW	13.44	41.64	0.96
NAB8	YW	13.33	41.8	0.66
NAMA	ZP	-7.51	31.04	1.56
NARO	XI	-1.07	35.87	1.92
NAZA	XI	8.57	39.29	1.73
NBI	AF	-1.27	36.8	1.71
NDEI	XI	-2.69	38.17	0.73
NEKE	XI	9.09	36.52	2.08
NGIT	UN	0.64	30.03	0.99
NJOM	ZP	-9.37	34.79	1.95

NNMO	YR	17.36	54.25	0.69
NURE	YJ	8.73	39.8	1.18
NYAN	UN	0.21	30.45	1.30
PAND	XD	-8.98	33.24	1.25
PIGI	ZP	0.23	32.32	1.25
PNDA	ZP	-6.35	31.06	1.07
POLI	YI	-9.77	33.87	0.47
PUGE	XD	-4.71	33.18	1.35
QALY	YR	12.7	53.49	0.04
QATE	ZF	9.38	41.47	2.15
QISH	XW	15.51	51.69	0.06
RAHO	YR	17.06	53.81	1.18
RAND	7C	11.85	42.66	0.88
RAYN	II	23.52	45.5	0.63
RODE	ZE	12.84	40.98	0.05
ROTI	ZP	1.63	33.6	1.11
RUGA	UN	-0.26	30.1	1.36
RUNG	XD	-6.94	33.52	1.23
RWEB	UN	0.32	30.49	1.28
SAAH	XW	15.57	48.86	0.78
SAHO	YR	17.11	54.68	1.18
SAKA	ZP	-0.32	31.74	1.26
SANA	XW	15.39	44.21	2.25
SAUM	XW	16.14	49.29	0.58
SAY	YR	15.35	44.2	2.25
SAYT	XW	15.22	51.25	0.06
SCH	YI	-10.18	34.03	0.52
SEHE	ZE	12.04	40.98	0.36
SEKE	ZF	12.62	39.03	2.26
SELA	XI	7.97	39.13	2.30
SEMP	UN	0.84	30.17	1.30
SENE	YJ	9.15	39.02	2.56
SEYU	XW	15.93	48.8	0.68
SHEE	YJ	10.00	39.89	1.30
SHIB	XW	15.5	43.91	2.63
SHIO	YR	17.19	54.17	0.55
SHUH	XW	15.61	50.92	0.26
SILE	II,ZE	12.41	41.19	0.48
SIMA	XW	17.56	52.32	0.66
SING	XD	-4.64	34.73	1.46

SMRE	ZF	13.20	39.21	1.98
SONG	ZP	- 10.67	35.65	1.12
SOOO	YR	17.08	54.88	0.14
SRDE	ZF	11.96	41.31	0.09
SUGH	XW	14.80	43.44	0.25
SULU	ZP	-4.57	30.09	0.09
SUMB	ZP	-7.95	31.62	1.84
TABU	XW	15.93	52.14	0.02
TALE	XI	0.98	34.98	1.82
TAMU	XW	17.29	49.93	0.67
TARA	XD	-3.89	36.02	1.27
TARI	XW	16.05	48.98	0.62
TAWI	XW	15.48	43.72	2.33
TEBE	AF	0.05	32.48	1.13
TEND	XI	11.79	41.00	0.42
TERC	XI	7.14	37.17	1.39
TEZI	AF	- 15.75	26.02	1.12
TINA	XW	16.53	52.08	0.20
TIOE	YR	14.67	40.87	0.04
TQHO	YR	17.06	54.43	0.04
TRUE	ZE,ZR	12.48	40.31	0.38
TUND	ZP	-9.30	32.77	1.66
U01	Z5	0.99	30.32	0.69
U02	Z5	-0.31	29.86	0.92
U03	Z5	-0.6	29.82	1.09
U04	Z5	0.34	30.04	1.64
U05	Z5	0.38	30.22	1.12
U06	Z5	0.19	30.45	1.27
U07	Z5	0.49	30.33	1.42
U08	Z5	-0.14	29.87	0.95
U09	Z5	0.02	30.08	0.91

U10	Z5	-0.1	30.38	1.41
U11	Z5	-0.44	30.56	1.45
U12	Z5	-0.53	30.16	1.65
U13	Z5	0.64	30.65	1.37
U14	Z5	0.71	30.06	0.95
U15	Z5	0.81	30.14	0.71
U16	Z5	0.69	30.28	1.56
U17	Z5	0.56	30.17	1.61
U18	Z5	0.73	30.37	1.56
U19	Z5	0.91	30.36	0.73
U20	Z5	0.03	29.77	1.24
U21	Z5	-0.28	30.04	1.00
U22	Z5	0.21	30.01	1.34
U23	Z5	0.35	29.89	4.45
UAYA	XW	15.71	42.69	0.01
URAM	XD	-5.09	32.08	1.12
UVZA	ZP	-5.10	30.39	0.99
WANE	XI	10.17	40.65	0.61
WASH	XI	8.99	40.17	0.83
WASH	XW	16.34	49.51	0.83
WELK	XI	8.29	37.78	1.90
WINO	ZP,ZE	-9.76	35.30	1.51
WLDE	ZF	11.82	39.59	1.88
WOLE	YJ	8.53	37.98	2.06
WUCE	ZF	11.51	39.61	1.91
YAF	YR	13.87	45.25	2.27
YAYE	ZF	11.86	38.00	2.63
YSLE	XW	14.94	44.28	2.56
ZOMB	AF	- 15.38	35.35	0.89
ZUWA	XW	15.73	43.02	0.09

Table S1. All seismic stations (station code, network code, latitude, longitude and elevation) used in the *S-SKS* tomographic inversion

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